

Analysis of Infrared Heating Rates in Observed Cloud Layers

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ABSTRACT

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A vertical heating rate profile is computed for an observed vertical distribution of cloud layers and liquid water content. One case looks at multiple-layered clouds in the western tropical Pacific Ocean and the other case examines a single-layer cloud in the eastern subtropical Atlantic Ocean. A broadband radiative transfer program is used to calculate heating and cooling rates. This method uses the Goody random statistical model to compute transmission through the clear air and cloud and assumes a linear absorption coefficient which is valid for small drops. Observations of the cloud scene are made from visible and infrared satellite imagery. The vertical distribution of clouds is determined from rawinsonde humidity profiles or ground based millimeter wavelength radar. Integrated cloud liquid water is computed from microwave brightness temperatures retrieved from satellite.

When the heating and cooling rates are integrated over 200 mb layers, cooling that results from the flux divergence at cloud top ranges from -0.28 °K/day to -0.63 °K/day, while the warming at the cloud base is less than 0.2 °K/day. This indicates that cloud top cooling due to infrared radiation plays a role in the diabatic heating and cooling of the atmosphere.

The cloud forcing in each case was also investigated. The multiple-layered clouds associated with the tropical convection increased cooling of the atmosphere at the cloud top, but the largest percent change was in the middle layers due to infrared heating. The single-layer cloud in the Atlantic Ocean case was too low to have much heating at the cloud base, but the cooling at cloud top was significant.

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CHAPTER 1

INTRODUCTION

Clouds modulate the overall diabatic heating of the atmosphere through release of latent heat, sensible heating and interaction with solar and terrestrial radiation. Latent heat is produced by phase changes in clouds and when precipitation is evaporated. Sensible heat is affected when clouds alter the temperature difference between the surface and the atmosphere. Clouds reflect solar radiation which cools the atmosphere and heat the atmosphere by emitting less longwave radiation to space than the surface. All of these processes impact atmospheric dynamics and thus are important in understanding the dynamics and lifetime of clouds.

Previous research has been conducted to determine qualitative estimates of the effects of clouds on diabatic heating by investigating the cloud diabatic forcing (CDF) [Siegmund, 1993]. Cloud forcing is the difference between diabatic heating in clear sky and cloudy conditions. Siegmund found CDF is dominated by release of latent heat in the storm track regions and the intertropical convergence zone.

On a local scale, differential radiative heating in clouds produces gradients of heating and cooling which influence atmospheric motions. Machado and Rossow [1993] found that the radiative effects of convective clouds may be realized more through atmospheric dynamics than through their direct effect on the earth's radiation balance. For example, radiative heating is found to be important to the dynamics of the upper-level

cloud shields of tropical cloud clusters [Houze, 1982]. Radiative heating in these cloud layers combined with mesoscale stratiform precipitation processes modifies the vertical heating profile sufficiently during development of tropical clusters to alter the large-scale vertical motion field. Ackerman *et al.*, [1988] modeled the radiative heating in cirrus clouds and found that destabilization occurs with an anvil of limited vertical extent and moderate ice water content (IWC). Such a heating rate gradient may induce dynamic responses that tend to lift and spread anvils.

The research conducted here uses observations from satellites to investigate the vertical profiles of longwave radiative heating and cooling rates in two different cloud scenarios. One case involves multiple layers of cumuliiform clouds over three sites in the tropical, western Pacific Ocean. The other case, in the subtropical region of the northern Atlantic Ocean, looks at the changes in heating rates over a one day for a single stratiform cloud layer.

Visible and infrared imagery from two geostationary satellites, the Japanese Geostationary Meteorological Satellite (GMS) and the European Meteosat, were used to identify the cloud scenes. Data collected by the Special Sensor Microwave/Imager on the United States Air Force Defense Meteorological Satellite Program (DMSP) satellites were used to determine the column integrated cloud liquid water. The method used, which was developed by Greenwald and Stephens [1995], simultaneously estimates the integrated cloud liquid water and water vapor amounts over the ocean.

The cloud profile for the tropical case is determined from rawinsonde observations using the refined relative humidity threshold method of Poore, Wang, and

Rossow [1994]. Millimeter wavelength radar was available for the Atlantic Ocean case to provide a more accurate location of the cloud layer.

A rough method is developed here for distributing the observed cloud liquid water among the layers. The method involves dividing the cloud into regions containing liquid water, mixed liquid/ice or ice based on temperature alone. Then the measured liquid water path is divided into the different regions based on cloud type and available water.

The heating and cooling rates are computed using a simple broadband radiative transfer program which utilizes the Goody transmission function. The observed vertical cloud boundaries and water and/or ice contents are input into the model. Finally, the heating and cooling profile is integrated into 200 mb layers to observe spatial and temporal variations in the infrared heating and cooling rates over four layers of the atmosphere.

CHAPTER 2

BACKGROUND

Radiation heating rates in clouds have been extensively analyzed through models [Fu and Liou, 1993; Slingo and Slingo, 1988; Starr and Cox, 1985; Stephens, 1978]. In this research, actual cloud scenes are investigated. This section describes how the cloud types are identified from satellite imagery and how the vertical distribution of these clouds can be determined from rawinsonde data. The simple radiative transfer program used to compute the heating and cooling rates is explained along with basic radiative transfer theory. Finally, the method used to determine cloud liquid water content from microwave length radiances is described.

2.1 Cloud Analysis

The locations and types of clouds can be identified through bispectral analysis of the visible and infrared imagery from geostationary meteorological satellites. The cloud base and vertical distribution cannot be determined from this data alone; however in the future, lidar or millimeter wavelength radar flown on satellites will provide such information. For this research, the vertical distribution of cloud is determined by locating the base and top of each cloud layer from rawinsonde humidity profiles [Wang and Rossow, 1995]. The top of the highest cloud layer is also computed by relating the retrieved satellite infrared brightness temperature to the rawinsonde temperature profile

and comparing it to the cloud top retrieved from the humidity profile. Where available, ceilometer and radar data provide additional information to locate the cloud layers.

2.1.1 Determination of cloud type

Geostationary meteorological satellites provide images of the earth disc every half hour or hour, depending on the satellite. Visible channels ($0.4\ \mu\text{m}$ to $0.7\ \mu\text{m}$) show clouds, which reflect the most visible light, as white and land, which reflects less visible light, as dark. Textural features like overshooting tops of cumulonimbus and the cellular features of stratocumulus can be identified due to the high spatial resolution ($1\ \text{km}$ - $4\ \text{km}$) of the visible channel ($0.4\ \mu\text{m}$ - $1\ \mu\text{m}$) and the contrast between the cloud and the earth's surface or areas of shadow. An infrared channel ($10\ \mu\text{m}$ - $12.5\ \mu\text{m}$) measures radiation that can be related to the temperature of the surface of cloud tops, land or ocean. A typical infrared image displays the cold pixels as white and the warm pixels as dark, so that clouds appear white on an image.

Cloud types can be determined using both the visible and infrared images according to the cloud classification scheme depicted in figure 2.1. Since temperature on average decreases with height in the troposphere, the infrared images are used to classify clouds as low, middle or high based on the temperature of the cloud top. Low clouds can be difficult to identify because the radiative temperature of the cloud is close to the temperature of the surface. High thin cirrus do not show up well in infrared images because the instrument detector sees the warmer earth's surface through gaps in the clouds. Thin cirrus may not even show up on visible images because they can be nearly transparent to visible light.

Visible and infrared images can be enhanced by assigning specific colors or gray shades to single pixel values or ranges of values to cause specific features of the cloud to stand out.

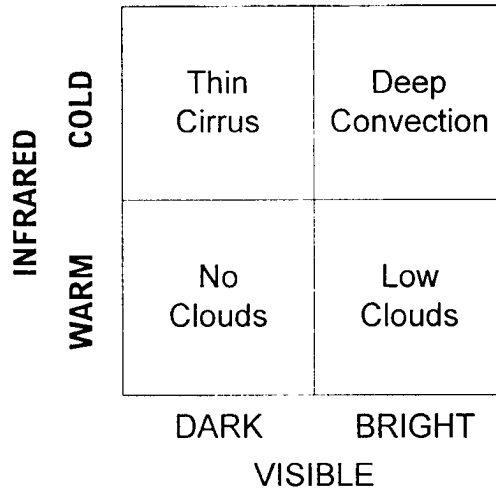


Figure 2.1: Schematic diagram of a bispectral cloud classification scheme [Kidder and Vonder Haar, 1995].

2.1.2 Distribution of cloud layers

Vertical profiles of temperature and humidity measured by rawinsondes provide a source of information on the vertical distribution of clouds [Wang and Rossow, 1995]. Rawinsondes measure temperature, relative humidity, wind speed and wind direction at various pressure levels from the surface to a maximum level. Relative humidity can provide some indication of the location of cloud layers because water vapor is saturated or supersaturated in a cloud. Use of the rawinsonde observations (RAOBS) to determine the vertical cloud distribution assumes that the sounding balloon ascended through the cloud and that the cloud is homogeneous. In addition, dewpoint measurements are usually

unreliable at temperatures below -40°C [Poore et al., 1994], so high cloud tops cannot be determined from this method.

The Air Weather Service [AWS, 1979], defines how the dewpoint depression, which is the difference between the dewpoint temperature and the air temperature, can be used to detect cloud base and top. The cloud base is indicated by a sudden decrease in dewpoint depression and the cloud top is indicated by a sharp increase. The dewpoint depression within the cloud must be smaller than

$$2^{\circ}\text{C for } T > 0^{\circ}\text{C}$$

$$4^{\circ}\text{C for } 0^{\circ}\text{C} < T < -20^{\circ}\text{C}$$

$$6^{\circ}\text{C for } T < -20^{\circ}\text{C}.$$

A more recent method developed by Poore et. al [1995] uses modified AWS thresholds, which are compared with surface observations as a quality check, to build a data set of layer cloud amount, type and height of cloud base and top.

Wang and Rossow [1995] revised the Poore et. al method to use only relative humidity measured from rawinsondes observations. The Wang and Rossow method establishes thresholds for the relative humidity of each level within the cloud and the increase/decrease in humidity at the lower/upper boundary of the cloud layer. These thresholds were determined by matching cases where clouds are observed in surface weather observations (SWOBS) to the relative humidity measured by RAOBS. Wang and Rossow identify cloud layers by examining the relative humidity at each level starting from the surface and increasing in height to the highest level of the profile.

Cloud base is determined when a level is encountered that meets one of the following conditions:

1. $RH \geq 87\%$
2. If the level is not the surface: $84\% \leq RH < 87\%$ and an increase of 3% from the layer below it.
3. If the level is the surface: $RH \geq 84\%$.

The levels above the base are temporarily considered to be cloudy as long as the relative humidity is at least 84%. When a level is encountered with lower relative humidity, the top of the cloud is sought by testing the levels downward from the highest cloudy layer against the following criteria:

1. $RH \geq 87\%$
2. If the level is not the top of the profile: $84\% \leq RH < 87\%$ and a decrease of 3% from the layer above it.
3. If the level is the top of the profile: $RH \geq 84\%$.

Finally, the moist layer is checked again and considered to be a true cloud only if the maximum relative humidity of the layer is greater than 87%. Additional cloud layers are detected by repeating the above process until the top of the profile is reached.

The relative humidity from RAOBS is with respect to liquid water at all temperatures and must be converted to a relative humidity with respect to ice at temperatures below 0°C to use the Wang and Rossow method because their thresholds are independent of temperature. Although the relative humidity of the mixed phase part of the cloud is actually between the relative humidity with respect to liquid and ice, the

magnitude of error in estimating relative humidity is countered by the systematic faults of the rawinsonde humidity element [Wang and Rossow, 1995].

The RAOBS obtained for this research report the humidity as a dewpoint temperature which must be converted into relative humidity with respect to liquid water and with respect to ice. Relative humidity is computed from the ratio of vapor pressure to saturation vapor pressure

$$RH = \frac{e}{e_s} \times 100\% \quad (2.1)$$

The saturation vapor pressure over liquid water, e_s , is computed from the formula [Bolton, 1980]

$$e_s(T) = 6.112 \exp\left(\frac{17.67T}{T + 243.5}\right) \quad (2.2)$$

where T is the ambient temperature in Kelvin. Equation (2.2) is also used to compute vapor pressure, e, where T is now equal to the dewpoint temperature.

When the ambient temperature is less than 0°C, the saturation vapor pressure over ice is computed from the formula [Emanuel, 1994]

$$e_{s_{ice}}(T) = \exp(23.33086 - (6111.72784 / T) + (0.15215 \ln(T))) \quad (2.3)$$

where T is the ambient temperature in Kelvin.

2.2 Calculation of heating and cooling rates

Heating and cooling rates are calculated using a 35-band infrared radiative transfer program with 10 mb resolution, referred to here as LWBAND. This section describes general radiative transfer theory in terms of how the LWBAND program computes the heating and cooling rates.

2.2.1 Radiative transfer theory

Radiation traveling through the atmosphere is attenuated due to absorption by atmospheric gases. The vertical distribution and spectral characteristics of the different atmospheric gases translates to a specific radiative heating and cooling profile. For example, water vapor has a strong cooling effect in the troposphere at infrared wavelengths. Overall, absorption of solar radiation by atmospheric gases generates a net heating which is approximately balanced by the cooling that is due to net infrared emission from the surface and atmosphere.

Reduction in the intensity of radiation is proportional to the amount of gas in a path through the atmosphere as described by Lambert's Law

$$dI_{\nu} = -k_{\nu} \rho_a I_{\nu} ds \quad (2.4)$$

where I_{ν} is the intensity at a specific frequency, k_{ν} is the mass absorption coefficient and ρ_a is the density of the absorbing gas. Integrating over a path length gives

$$I_{\nu}(s_2) = I_{\nu}(s_1) \exp\left[-\int_{s_1}^{s_2} k_{\nu} \rho_a ds\right] \quad (2.5)$$

The attenuation of intensity over the path is represented by the optical path or thickness and is defined as

$$\tau(s_1, s_2) = \int_{s_1}^{s_2} k_v \rho_a ds \quad (2.6)$$

By assuming the atmosphere is made up of plane parallel layers, the normal optical path can be described as

$$\tau(s) = \frac{\tau(z)}{\mu} \quad (2.7)$$

where $\mu = \cos\theta$ and θ is the zenith angle.

Equation 2.5 can then be expressed as

$$I_v(z_2) = I_v(z_1)e^{-\tau(z)/\mu} \quad (2.8)$$

Radiant intensity is decreased by the exponential function referred to as monochromatic transmittance

$$T_v = e^{-\tau(z)/\mu} \quad (2.9)$$

The upward and downward intensities of radiation, assuming thermodynamic equilibrium and a plane-parallel atmosphere, are expressed as

$$I_v^+(\tau, +\mu) = B_v(T_s)e^{-(\tau' - \tau)/\mu} - \int_{\tau}^{\tau'} B_v[T(\tau')]e^{-(\tau' - \tau)/\mu} \frac{d\tau'}{\mu} \quad (2.10)$$

$$I_v^-(\tau, -\mu) = I_v(0, \mu)e^{-\tau/\mu} + \int_0^{\tau} B_v[T(\tau')]e^{-(\tau - \tau')/\mu} \frac{d\tau'}{\mu} \quad (2.11)$$

The upwelling intensity described in equation 2.10 consists of two terms which are emission from the earth's surface and emission from the atmosphere. The downwelling intensity in equation 2.11 is due to emission from the cosmos and the atmosphere; however emission from the cosmos is usually assumed to be zero. The Planck function,

$B_\nu(T)$, describes the intensity from the surface and atmosphere at a particular frequency.

Optical depth is defined to be zero at the top of the atmosphere and τ' at the earth's surface.

Flux is the sum of the directional intensities

$$F = 2\pi \int_0^1 I_\nu d\mu \quad (2.12)$$

Using equations 2.10 and 2.11, the monochromatic upward and downward fluxes can then be expressed as

$$F_\nu^+(\tau) = 2\pi B_\nu(T_s) \int_0^1 e^{-(\tau'-\tau)/\mu} \mu d\mu + 2 \int_0^1 \int_\tau^{\tau'} \pi B_\nu [T(\tau')] e^{-(\tau'-\tau)/\mu} d\tau' d\mu \quad (2.13)$$

$$F_\nu^-(\tau) = 2 \int_0^1 \int_0^\tau \pi B_\nu [T(\tau')] e^{-(\tau-\tau')/\mu} d\tau' d\mu \quad (2.14)$$

In order to analytically compute these fluxes, a diffuse transmittance is defined as

$$T_\nu^f(\tau) = 2 \int_0^1 T_\nu(\tau/\mu) \mu d\mu = 2 \int_0^1 e^{-\tau/\mu} \mu d\mu \quad (2.15)$$

The flux transmission can be approximated by using transmission along the direction defined by the zenith angle $\theta = \cos^{-1}(1/\beta)$ where the diffusivity factor $\beta = 1.66$. The flux transmission is then written in terms of optical depth as

$$T_\nu^f = e^{-\beta\tau} \quad (2.16)$$

Integrating the monochromatic flux over all frequencies in the thermal IR spectrum gives

$$F^\pm(z) = \int_0^\infty F_\nu^\pm(z) d\nu \quad (2.17)$$

Net flux is defined as upward flux (F^+) minus downward flux (F^-), and the change in flux for a layer of atmosphere is

$$\Delta F = F(z + \Delta z) - F(z) \quad (2.18)$$

The divergence of net fluxes over a layer of the atmosphere determines whether heating or cooling will occur in that layer. A layer of air is heated if the net flux density at the top of the layer ($z+\Delta z$) is smaller than the bottom (z).

The heating rate is proportional to the local rate of change of flux by

$$\frac{\partial T}{\partial t} = -\frac{1}{\rho C_p} \frac{dF(z)}{dz} \quad (2.19)$$

where C_p is the specific heat at constant pressure, ρ is air density. As seen in the above equation, a positive heating rate occurs when change in net flux with height is negative. A negative heating rate, or cooling, results when the slope of the net flux is positive.

This project only investigates heating rates due to infrared radiation, so only wavelengths from 4.5 μm to 250 μm are considered. In this region of the spectrum, the greatest contribution to cloud radiative heating and cooling occurs in the atmospheric window (8 μm to 12 μm) because the atmosphere is more transparent to radiation at these wavelengths, allowing radiation to easily reach the cloud base from the surface. The rotational and vibrational bands, 13.9 μm - 250 μm and 4.55 μm - 8.33 μm respectively, contribute less since radiation at these wavelengths originates from layers near the cloud where the temperature is nearly the same, thus resulting in a smaller flux divergence.

2.2.2 Radiative transfer with clouds

Clouds emit infrared radiation and also reflect and transmit the infrared radiation emitted from the atmosphere and surface of the earth. Heating occurs at the base of a cloud due to absorption of infrared radiation emitted by the surface of the earth. Cooling occurs at the top of the cloud because the cloud emits to space at a lower temperature.

The amount of heating and cooling is dependent on the height and optical depth of the cloud. Optical depth for a cloud is defined by

$$\tau_{cld} = \int \sigma_{ext} dz \quad (2.20)$$

where σ_{ext} is the volume extinction coefficient. The extinction coefficient consists of absorption and scattering ($\sigma_{ext} = \sigma_{abs} + \sigma_{sca}$), but since scattering by cloud particles at infrared wavelengths is small, only absorption will be considered here. The absorption coefficient for spherical droplets is a function of droplet size distribution and absorption efficiency, Q_{abs} , and is defined as

$$\sigma_{abs} = \int Q_{abs}(v, r) \pi r^2 n(r) dr \quad (2.21)$$

Absorption efficiency is computed from Mie theory but is found to increase linearly with particle radius for small drops at long wavelengths as shown in figure 2.2.

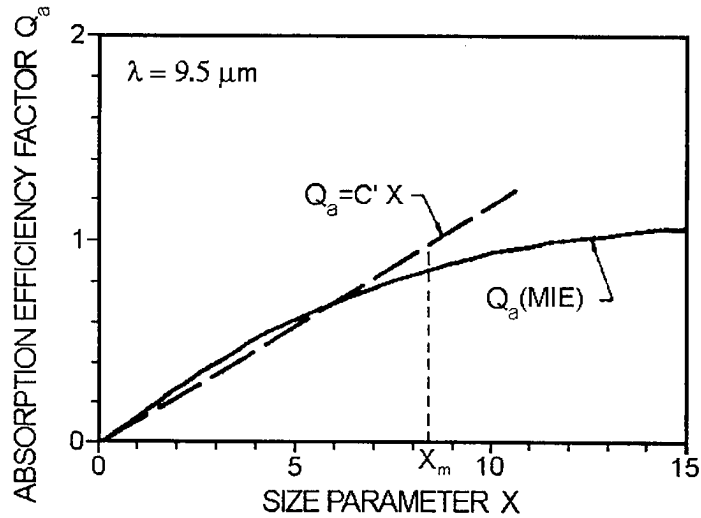


Figure 2.2: Absorption efficiency factor, Q_a , as a function of size parameter, $x (=2\pi r/\lambda)$ [from Pinnick et al. 1979].

This linear relation is written as

$$Q_{abs}(\nu, r) = C(\nu)r \quad (2.22)$$

and is valid for clouds with droplets of radius smaller than a maximum radius defined by maximum size parameter $x_m = 2\pi r_m / \lambda$. For this project the linear coefficient, $C(\nu)$, was computed by finding the value of $C(\nu)$ that gives the maximum radius with a 10% error between Q_{abs} computed from Mie theory and $C(\nu) \times r$.

The optical thickness of the cloud can now be written as

$$\tau_{cloud} = \int C(\nu) \left\{ \int \pi r^3 n(r) dr \right\} dz \quad (2.23)$$

The volume absorption coefficient, σ_{abs} , is converted to a mass absorption coefficient by dividing by liquid water content, w , so

$$k_{abs} = \frac{\sigma_{abs}}{w} \quad (2.24)$$

Liquid water content, which has units of g/m^3 , is a function of the droplet size distribution and is defined as

$$w = \int \frac{4}{3} \pi r^3 \rho_l n(r) dr \quad (2.25)$$

Substituting in for σ_{abs} and w in equation 2.24 ,

$$k_{abs,\nu} = \frac{C(\nu) \int \pi r^3 n(r) dr}{\frac{4}{3} \rho_l \int \pi r^3 n(r) dr} = \frac{3}{4} \frac{C(\nu)}{\rho_l} \quad (2.26)$$

Note that the mass absorption coefficient is now independent of droplet distribution.

Using the mass absorption coefficient, the optical depth of a cloud in the infrared spectrum can be approximated by

$$\tau_{cld} \approx k_{abs,v} w \Delta z \quad (2.27)$$

2.2.3 Calculation of transmission

In LWBAND, diffuse transmission for molecular absorption is computed using the Goody random statistical model. Calculation of the transmission functions for water vapor, carbon dioxide and ozone for a given spectral interval $\Delta\nu$ is through the equation

$$T_v^f = \exp \left[\frac{-\sigma \beta u \delta}{[1 + \sigma \beta u \pi \alpha_l]^{1/2}} \right] \quad (2.28)$$

where β is the diffusivity factor of 1.66, and σ (mean line intensity), δ (mean line spacing) and α_l (half width) are all tabulated parameters. The path length, u , is defined by

$$u = \int_{z_1}^{z_2} \rho_a dz = \int_{p_1}^{p_2} \rho_a \left(-\frac{1}{\rho g} dp \right) \quad (2.29)$$

Substituting for mixing ratio, $r = \rho_a/\rho$, and integrating from the surface ($p_1 = p_s$) to the top of the atmosphere ($p_2 = 0$) results in

$$u(0, p_s) = \frac{r p_s}{g} \quad (2.30)$$

The path length for a layer is

$$u(l) = \frac{\bar{r} \Delta p(l)}{g} \quad (2.31)$$

where \bar{r} is the mean mixing ratio for the layer of the particular absorbing species.

The Van de Hulst-Curtis-Godson (VCG) approximation is used to correct for atmospheric inhomogeneities by scaling the path u and pressure as follows

$$\tilde{p}\tilde{u} = \int p du \quad (2.32)$$

$$\tilde{u} = \int du \quad (2.33)$$

then,

$$\tilde{p} = \frac{\int p du}{\int du} \quad (2.34)$$

The half width, α_1 , in equation 2.28 is multiplied by $p' = \tilde{p}/p_s$ and the path length, u , is replaced with \tilde{u} .

Transmission for the water vapor continuum is computed using an absorption coefficient calculated from [Roberts, et al., 1976]

$$k_v = k_{0,v} \exp\left[1800\left(\frac{1}{T} - \frac{1}{296}\right)\right] \quad (2.35)$$

where $k_{0,v}$ (cm^2/g) is

$$k_{0,v} = 4.18 + 5578 \exp(-0.00787v) \quad (2.36)$$

The diffuse transmission for the water vapor continuum is then computed using equation 2.16.

Diffuse transmission for a cloud is computed from

$$T_{cld}^f = e^{-\beta\tau_{cld}} \quad (2.37)$$

Transmission for a particular band and atmospheric level is equal to the product of the transmission of all absorbing species in the band.

$$T = T_{H_2O} \times T_{CO_2} \times T_{O_3} \times T_{cont} \times T_{cld} \quad (2.38)$$

2.2.4 Comparison of LWBAND output to CAGEX data

Data from the CERES/ARM/GEWEX Exercise (CAGEX) was used in LWBAND to show how the model performs. The CAGEX data is a public access set of input data, calculated fluxes and validating measurements over the ARM CART site in Oklahoma. Upward and downward flux profiles are provided from the δ -four stream shortwave and longwave radiative transfer code of Fu and Liou [1993]. The data includes observations of cloud base and thickness and sounding data. Liquid water content is not provided but can be computed from measured emissivity or cloud optical depth.

Clear sky computations were made with LWBAND using the CAGEX sounding data. The slope of the net flux, which is proportional to the heating and cooling rate by a factor of $1/\rho C_p$ is shown in figure 2.3. Considering the simplicity of the LWBAND model, there is a fair comparison between the LWBAND and CAGEX results.

The heating rates for a cloudy case were computed using observations of cloud base and thickness directly from the CAGEX sounding information and computing liquid water content from measured optical depth. Figure 2.4 shows the comparison of CAGEX and LWBAND net flux slopes with a cloud layer present. The 300 meter thick cloud is based at 720 mb with a liquid water path of 25 g/m².

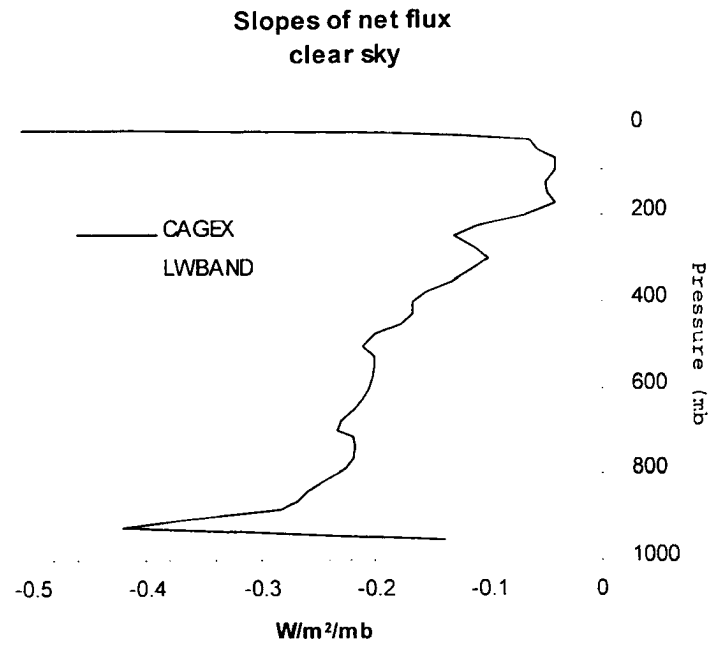


Figure 2.3: Comparison of CAGEX and LWBAND net flux slope in clear sky.

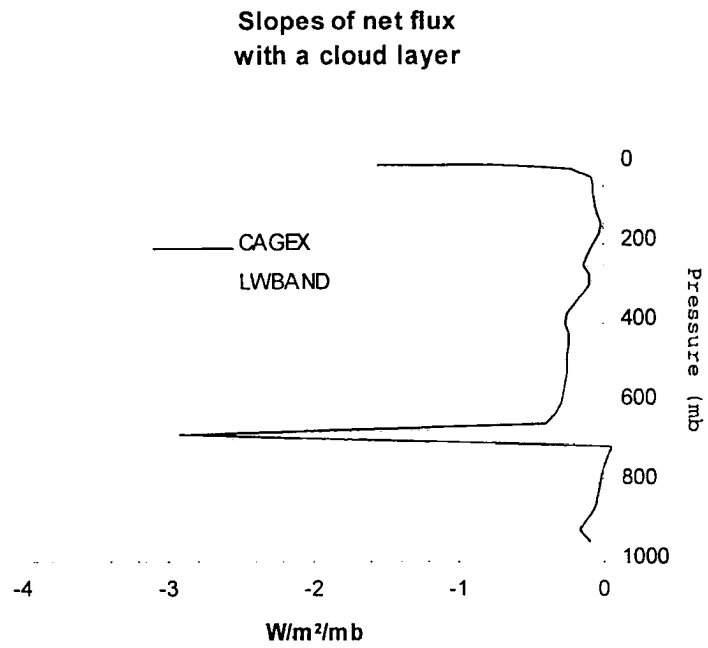


Figure 2.4: Comparison of CAGEX and LWBAND net flux slopes with a cloud layer present.

2.2.5 Sensitivity of heating rate to water content and cloud distribution

The magnitude of infrared heating and cooling is directly related to cloud height, thickness and the amount and distribution of liquid water in the cloud. The effects of these parameters on heating rates are shown in the following sensitivity study. The LWBAND radiative transfer model was used to compute the heating rates, and a standard, mid-latitude, summer McClatchey sounding provided the temperature and moisture profiles.

Figure 2.5 shows the effect of cloud height on heating and cooling rates. A single 300 meter thick cloud is raised from a cloud base of 1 km to 9 km. In Case I the cloud has a LWC of 0.2 g/m^3 and in Case II the cloud has a LWC of 0.6 g/m^3 . There is strong cooling at cloud top and heating at cloud base for both cases. The cooling rate increases with increasing height of the cloud up to 7 km for Case I and 4 km for Case II and then begins to decrease. The increase in heating rate results from the larger amount of infrared flux from the atmosphere below the cloud, the greater difference in temperature between the cloud base and the surface, and the decrease in air density.

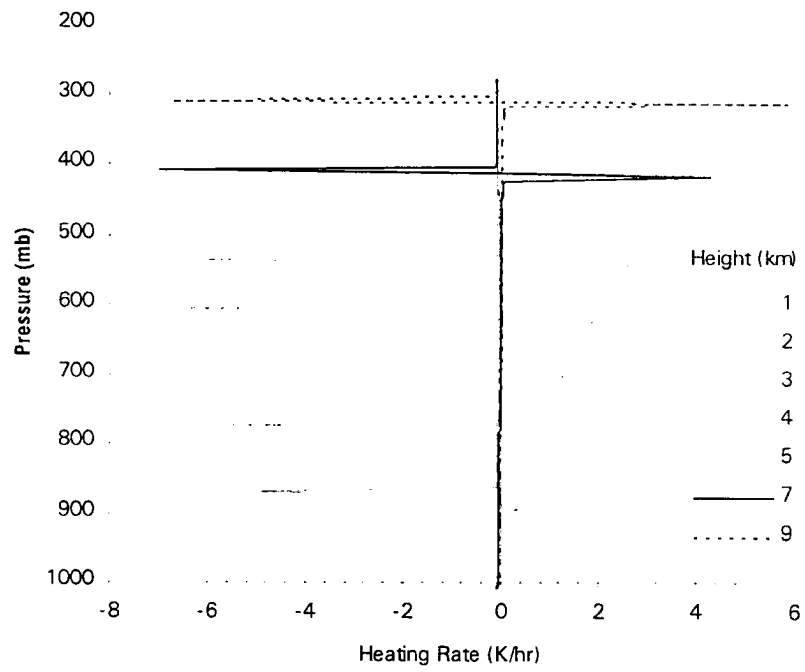
The thickness of a cloud also affects the heating rate since transmission through a cloud depends on optical depth (equation 2.37) which is proportional to the thickness of the layer (equation 2.27). Figure 2.6 shows the heating and cooling rates for a cloud with the base fixed at 1.5 km (850 mb) while the thickness of the cloud increases from 200 meters to 3000 meters. The cloud has a LWC of 0.2 g/m^3 . The cooling rate at the top of the cloud increases with increasing thickness but there is no change in heating rate.

The sensitivity of the heating and cooling rate to changing LWC is shown in figure 2.7. The cloud layer is fixed with a base at 1.5 km (850 mb) and a top at 1.8 km

(820 mb). The liquid water is distributed evenly throughout the cloud and varies from 0.01 g/m^3 to 1.5 g/m^3 . The heating and cooling rates both increase with increasing LWC due to an increase in the cloud optical depth. Figure 2.8 shows the peak heating and cooling values for each LWC. From this plot, both curves approach a constant value when LWC exceeds 1.0 g/m^3 . The heating rate approaches a maximum of 0.9 K/hr and the cooling rate approaches a minimum of -7.0 K/hr .

The effect of the distribution of cloud water on heating and cooling is investigated next. The cloud for this simulation is 300 meters thick cloud with a base at 1.5 km and an equivalent liquid water path of 30 g/m^2 . The liquid water is distributed so that the maximum water is at the top, bottom or middle of the cloud or evenly distributed as shown in figure 2.9. Figure 2.10 shows the heating and cooling profiles. The strongest cooling occurs when the largest amount of water is located at cloud top (distribution A). The next highest cooling occurs when the liquid water is evenly distributed (distribution D). When the maximum amount of water is at the base of the cloud (distribution B), the cooling rate is lowest and peaks below the cloud top. The peak cooling rate also occurs below the cloud top when the maximum amount of water is in the middle of the cloud (distribution C). Distributions B and D cause the highest heating rates which occur nearer to the base of the cloud. The heating that results from distributions A and C is stretched out over a larger portion of the cloud. Heating occurs at the highest level in the cloud when the maximum amount of water is at the top of the cloud.

I.



II.

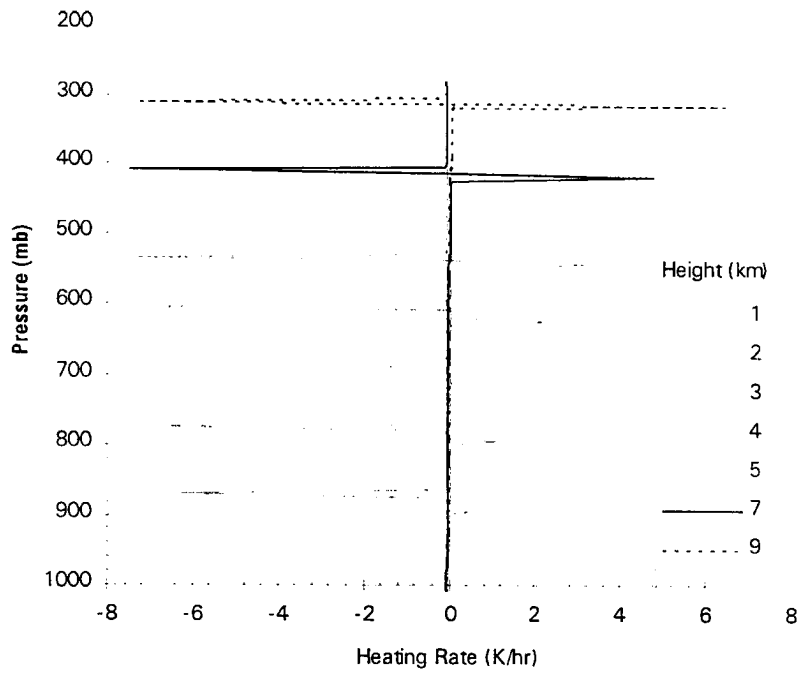


Figure 2.5: Change in heating rates with height for a 300 meter thick cloud with a LWC of 0.2 g/m³ (Case I) and 0.6 g/m³ Case II).

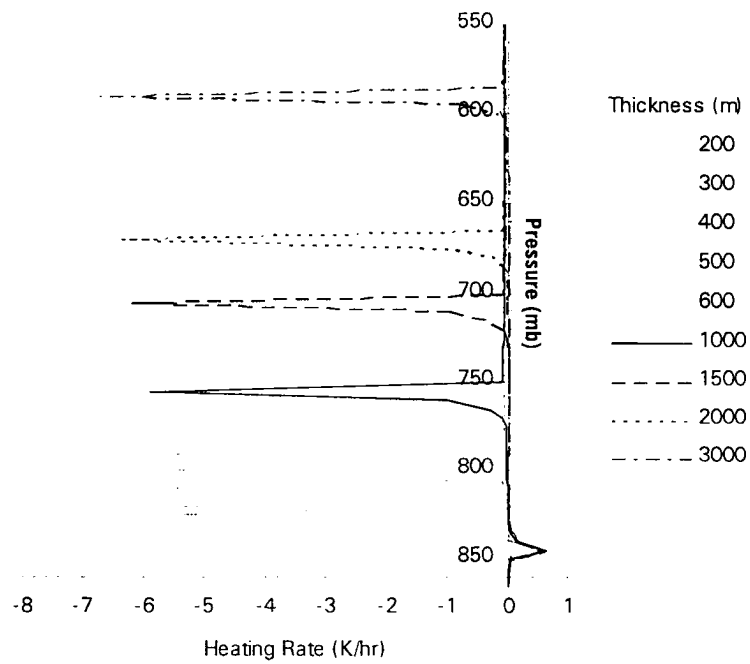


Figure 2.6: Heating and cooling rates when the thickness of the cloud is changed. Cloud base is held fixed at 850 mb and $LWC = 0.2 \text{ g/m}^3$.

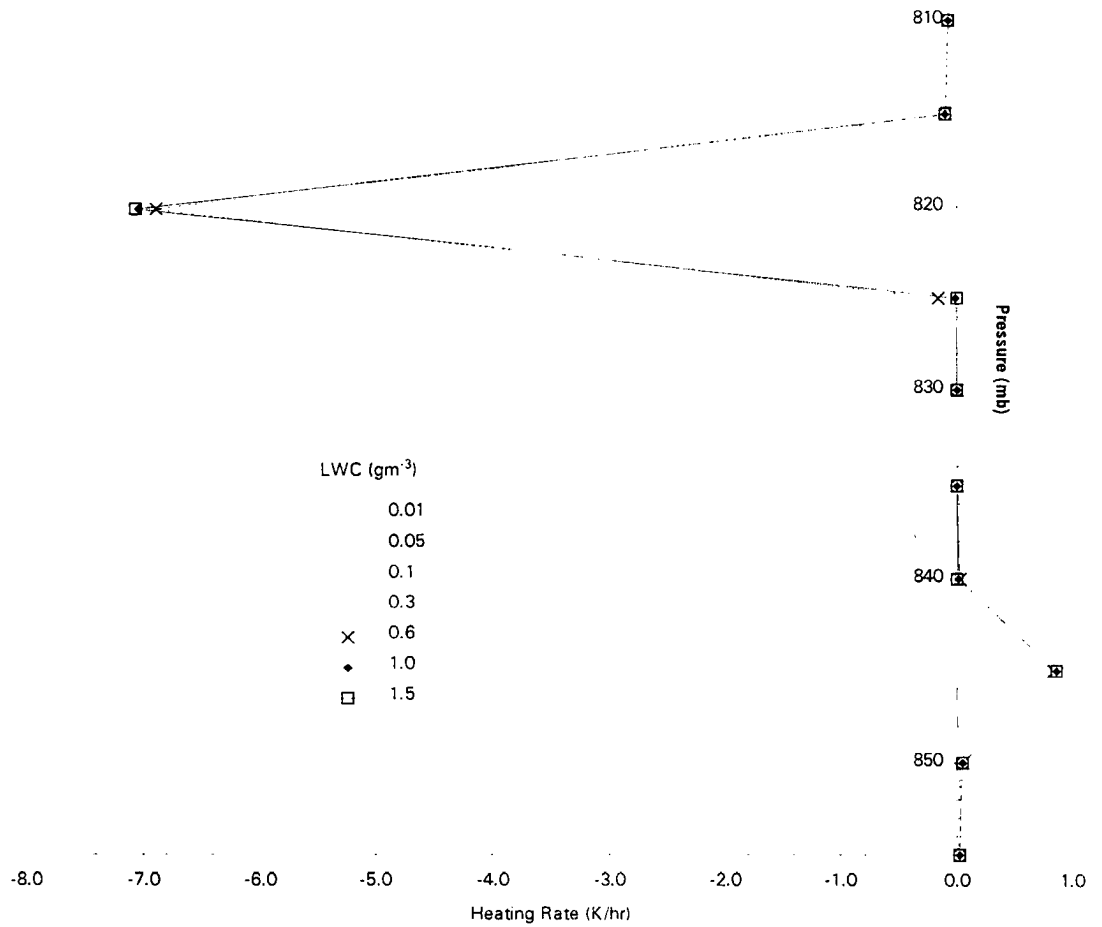


Figure 2.7: Heating rates for a range of LWC for a cloud with a base at 850 mb and a top at 820 mb.

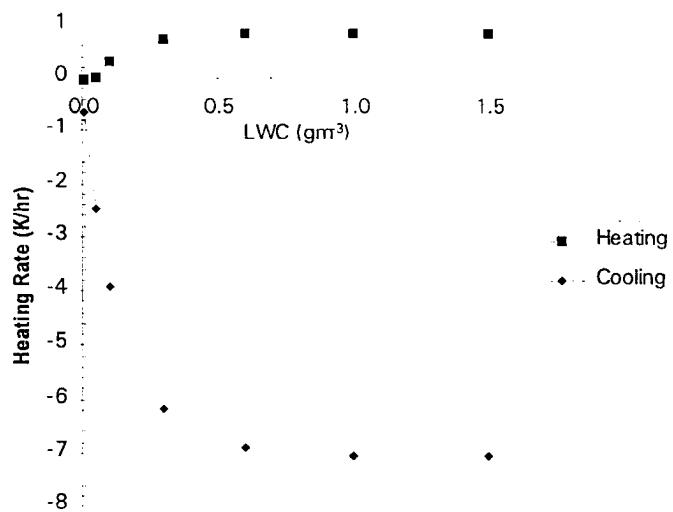


Figure 2.8: Trends of maximum heating and cooling rates over a range of LWC.

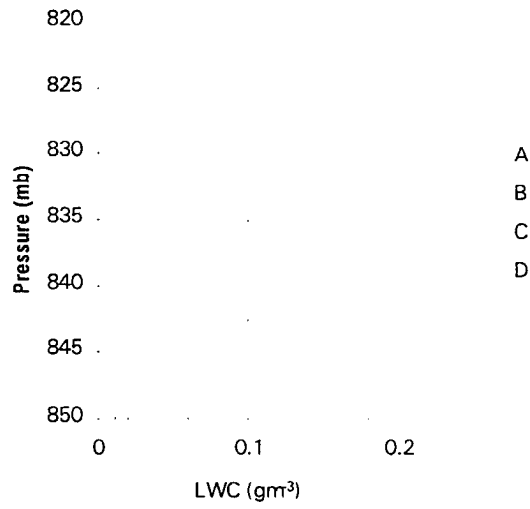


Figure 2.9: Different distributions of cloud liquid water for a 300 meter thick cloud with a liquid water path of 30 g/m².

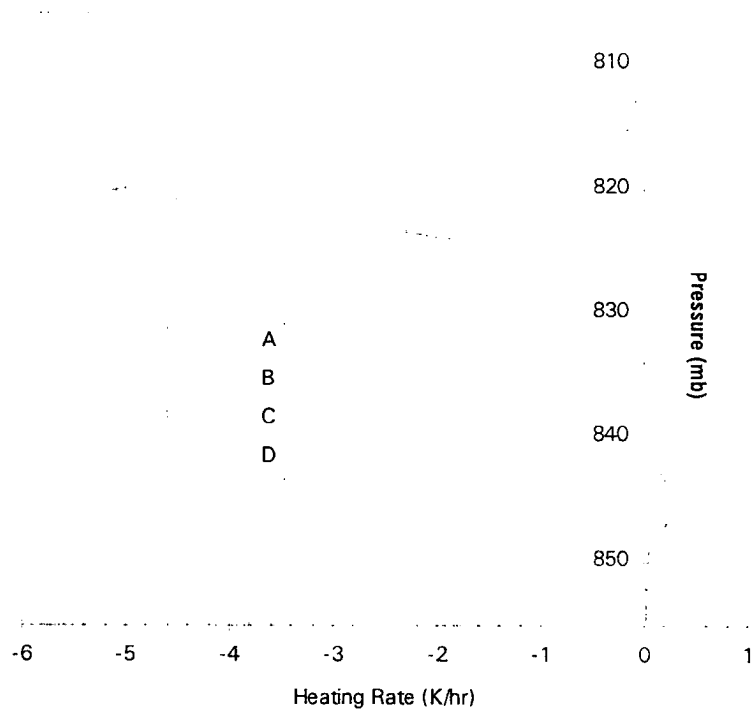


Figure 2.10: Heating rates for the distributions of water in figure 2.9.

2.3 Liquid water retrieval using passive microwave

2.3.1 Theory

The microwave spectrum is defined from 3 GHz (10 cm) to 1000 GHz (0.3 mm). There are absorption lines for water vapor and molecular oxygen in this range which make microwave remote sensing useful for measuring atmospheric water content; however, a problem with using microwave sounding is that emissivity values of the earth's surface vary from 0.4 to 1.0 in the microwave region. Most microwave retrieval methods to determine water vapor and cloud liquid water can only be used over the ocean where the emissivity typically only ranges between 0.4 and 0.5, although a method that uses a combination of IR and microwave data has been developed to determine cloud liquid water over land [Jones and Vonder Haar, 1990].

The ocean appears cold at microwave frequencies but the overlying atmosphere and cloud layer increases the emitted radiation as viewed by a satellite. Column integrated water vapor (called precipitable water) and cloud liquid water can be determined by measuring the change in the emission.

The presence of ice or precipitation can alter the retrieved microwave radiance. Ice scatters radiation at microwave frequencies if the ice particles are large compared to the wavelength of the incident radiation. Raindrops and cloud drops larger than 100 greatly absorb and scatter microwave radiation. Therefore, retrieval of cloud liquid water is not as accurate for precipitating clouds and high clouds, such as cirrus, that contain ice.

2.3.2 Description of method

The total amount of water contained in an atmospheric column is determined from a physical retrieval method [Greenwald and Stephens, 1995] that uses passive microwave measurements of brightness temperatures. This method, hereafter referred to as the Greenwald method, employs simple, vertically integrated forms of the radiative transfer equation with minimal empiricism that is used only for calibration purposes. The method simultaneously computes the liquid water path and water vapor content for a non-precipitating, ice-free, oceanic region given microwave vertical and horizontally polarized brightness temperatures, infrared brightness temperature and sea surface temperature.

The brightness temperature measured by a microwave radiometer is a function of atmospheric transmittance and surface emissivity. In the microwave region, Planck radiance is directly proportional to temperature, which is known as the Rayleigh-Jeans distribution. Thus the radiative transfer equation can be written as

$$T_B(z_i) = \epsilon_s T_s \tau(z_0, z_i) + \int_{z_0}^{z_i} T(z') \frac{d\tau(z', z_i)}{dz'} dz' + R_s \tau(z_0, z_i) \int_{z_i}^{z_0} T(z') \frac{d\tau(z_0, z')}{dz'} dz' \quad (2.40)$$

The first term is the emission from the surface, the second term is the integrated atmospheric emission, and the last term is the downwelling radiation emitted by the atmosphere that is reflected by the surface and retransmitted to the satellite. The equation contains the atmospheric transmission function, τ , polarized surface emissivity, ϵ_s , polarized surface reflectivity, R_s or $(1 - \epsilon_s)$, and surface temperature, T_s .

In the microwave spectrum, the surface emissivity varies considerably over land depending on the moisture content of the soil and type of vegetation. Over the ocean, the surface emissivity varies with surface roughness, salinity and amount of sea foam. The Greenwald method currently is valid only over the ocean. The surface emissivity is calculated from existing models that use microwave brightness temperatures and sea surface temperature as input.

Assuming an isothermal atmosphere and absorption by water vapor in the boundary layer only, equation (2.40) can be simplified to

$$T_b(z_r) = T_s[1 - \tau^2(1 - \epsilon_s)] \quad (2.41)$$

The square of the atmospheric transmission function contains the transmission factors for oxygen, cloud liquid water and water vapor.

$$\tau^2 = \tau_w^2 \tau_l^2 \tau_o^2 \quad (2.42)$$

The transmission factor for water vapor is a function of the water vapor absorption coefficient, viewing angle and water vapor content.

$$\tau_w = \exp\left(-\frac{\kappa_w W}{\mu}\right) \quad (2.43)$$

The transmission factor for cloud liquid water is a function of the absorption coefficient for liquid water, viewing angle, and liquid water path.

$$\tau_l = \exp\left(-\frac{\kappa_l L}{\mu}\right) \quad (2.44)$$

The transmission factor for oxygen depends on temperature and is determined by the empirical equation

$$\tau_{ox} = a + bT_a + cT_a^2 + dT_a^3 \quad (2.45)$$

To solve for W and L, equation (2.41) is first used to obtain the horizontally and vertically polarized brightness temperatures. The difference between the brightness temperatures is

$$\Delta T_B = T_s (\epsilon^H - \epsilon^V) \tau_w^2 \tau_L^2 \tau_{ox}^2 \quad (2.46)$$

Substituting equations (2.43) and (2.54) into equation (2.46) and rearranging

$$\kappa_{\nu} L + \kappa_w W = -\frac{\mu}{2} \ln \left(\frac{\Delta T_B}{T_s (\epsilon^H - \epsilon^V) \tau_{ox}^2} \right) \quad (2.47)$$

Equation (2.47) is then applied to two microwave channels, such as the 19 GHz and 37 GHz channels, to provide a set of linear equations

$$W = \frac{\tau_1 \kappa_{\nu 37} - \tau_2 \kappa_{\nu 19}}{\delta} \quad (2.48)$$

$$L = \frac{\tau_2 \kappa_{w19} - \tau_1 \kappa_{w37}}{\delta} \quad (2.49)$$

where,

$$\delta = \kappa_{w19} \kappa_{\nu 37} - \kappa_{w37} \kappa_{\nu 19} \quad (2.50)$$

$$\tau_1 = -\frac{\mu}{2} \ln \left[\frac{\bar{T}_{19} - T_{19v}}{T_s (1 - \epsilon_{19v}) \tau_{ox19}^2} \right] \quad (2.51)$$

$$\tau_2 = -\frac{\mu}{2} \ln \left[\frac{\bar{T}_{37} - T_{37v}}{T_s (1 - \epsilon_{37v}) \tau_{ox37}^2} \right] \quad (2.52)$$

The τ_1 and τ_2 physically represent optical depths of the atmospheric water in non-precipitating regions and excluding ice. The \bar{T}_{37} and \bar{T}_{19} values compensate for the vertical change in water vapor and temperature and other unmodelled effects.

CHAPTER 3

DATA

3.1 Case selection

Two areas of study were selected for this research, one in the equatorial warm pool of the western Pacific Ocean and one in the subtropic region of the eastern Atlantic Ocean. Marine cases were necessary because the Greenwald method [Greenwald, 1995] for calculating liquid water is valid only over the ocean. Data for the first case were available for multiple sites which allows for the investigation into the spatial heating rate gradients. The second case consists of only one site, so heating rates were analyzed over the course of one day.

3.1.1 Case I

The first case was selected with the intent of using data from the Tropical Ocean and Global Atmosphere Coupled Ocean-Atmosphere Response Experiment (TOGA COARE) because of the extensive data collected during this field experiment. TOGA COARE was conducted in the region of 145°E to 155°E and 5°S to 5°N from November 1992 to February 1993 to study ocean-atmospheric processes in the western tropical pacific. TOGA COARE data sets used here were the 6 hourly balloon soundings from islands and research ships and half hourly visible and infrared images from the GMS satellite. Figure 3.1 shows a map of this region with rawinsonde launch sites identified.

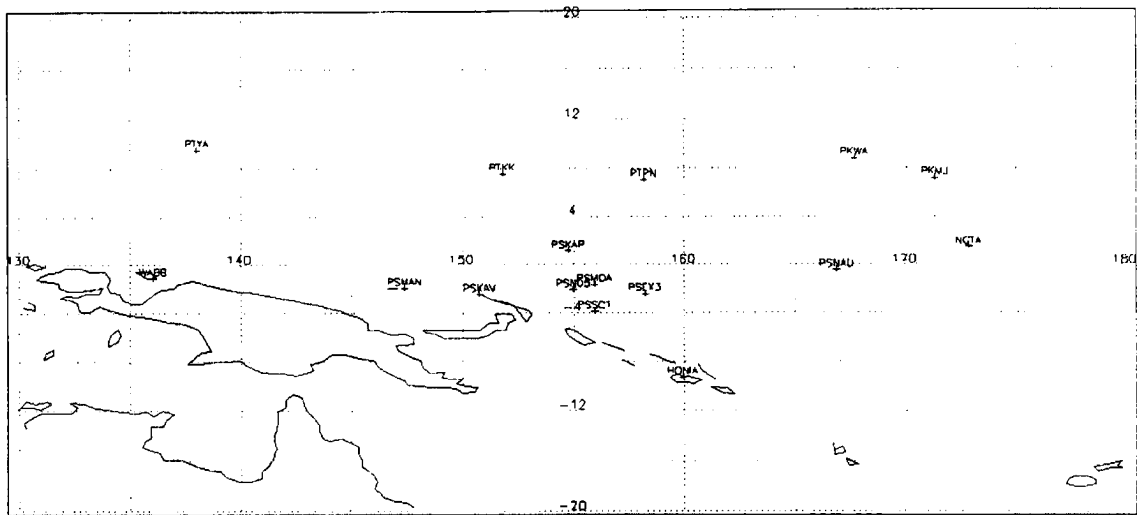


Figure 3.1: Location of rawinsonde sounding sites in the TOGA COARE region.

Calculation of radiative heating rate requires knowledge of the distribution of cloud layers and their liquid water content as well as the vertical temperature, pressure and humidity profiles. Since SSM/I is used to measure the liquid water path, the primary cases for analysis must coincide with DMSP overflights of the region. The selected SSM/I data set consists of microwave data collected on J356 from 2303Z to 2310Z. The nearest GMS IR and visible images are from J356 at 2315Z. Analysis is further limited to the areas immediately surrounding sites where rawinsondes were launched because the soundings are the only means of determining the vertical cloud structure in this case. As shown in figure 3.2, the closest sounding data are available for J357 at 0000Z, approximately one hour later.

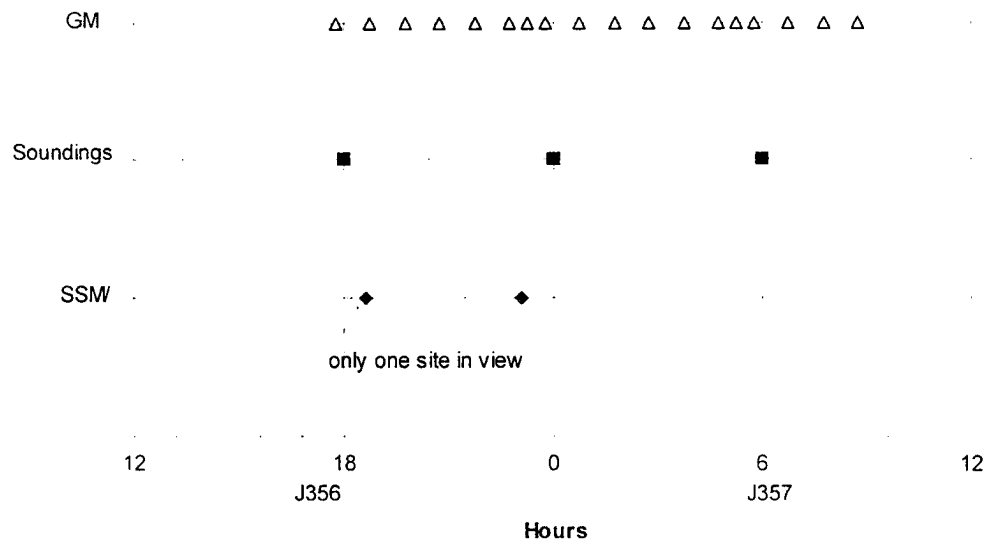


Figure 3.2: Timeline of the satellite and rawinsonde data retrieval times for December 21-22, 1992 (J356-J357).

The primary analysis time for which the heating and cooling rates are computed is established as 0000Z (10:00 am, local time) on J356. Heating and cooling rates are also computed six hours earlier and six hours later using the rawinsonde and IR data to determine the vertical cloud profile but the same liquid water amount from 0000Z because SSM/I data is not available. The exception is that SSM/I data is available only for PSMOA at 1800Z.

Three sounding sites were selected based on their proximity to each other and because the rawinsonde data were available every 6 hours. The selected sounding sites are: the island of Kavieng (PSKAV) and the research ships Xiangyanghong 5 (PSN05) and Moana Wave (PSMOA).

The resolution of the data is limited by the SSM/I 37 GHz channel which has a resolution of 37 km x 29 km. The location of each sounding site is shown relative to the nine nearest data points in figure 3.3.

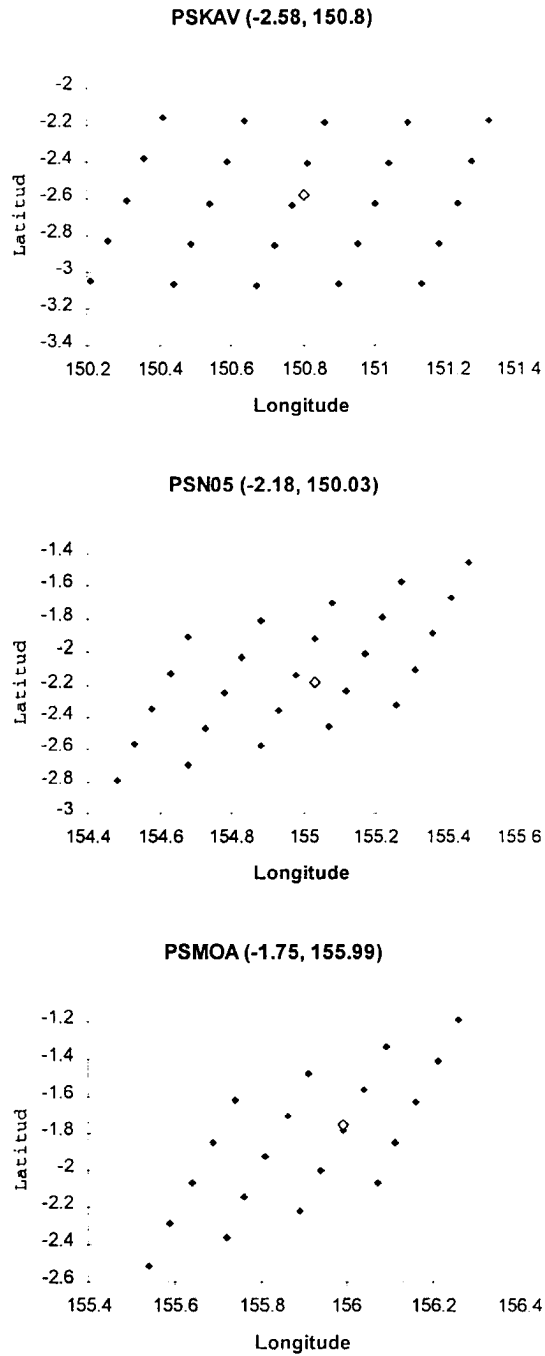


Figure 3.3: Locations of sounding sites relative to gridded data points.

3.1.2 Case II

Case II looks at a single cloud layer in a subtropical marine region. The area selected was studied in the ASTEX (Atlantic Stratocumulus Transition Experiment) phase of FIRE (First ISCCP Regional Experiment). This field experiment occurred during the month of June 1992 on Porto Santo, an island in the Madeira Islands located at 33.08 °N latitude and 16.35 °W longitude in the north Atlantic Ocean. A map of the island and location of the experiment site is shown in Figure 3.4.

Figure 3.4: Map of the ASTEX region [from Huebert, et al., 1996].

ASTEX was chosen because of the high occurrence of single layer stratus clouds that transition to stratocumulus and because of the extensive ensemble of in situ and

remote sensing instruments. A description of the instruments and data sets can be found in Cox *et al.* [1993].

The surface observation and remote sensing data sets acquired from ASTEX include rawinsonde, ceilometer, microwave radiometer, and Doppler 8.7 mm radar. Satellite data consists of IR and visible images from the European Meteosat-3 satellite and microwave radiometer data from SSM/I on the DMSP satellite. The availability of these data sets for June 14 (J166) is shown in figure 3.5. June 14 was selected for analysis based on ground observation reports of overcast, non-precipitating conditions.

Cloud liquid water path and infrared data are gridded based on the resolution of the 37 Ghz channel on the SSM/I instrument which is 37 km x 29 km. Figure 3.6 shows the proximity of these gridded data points to Porto Santo.

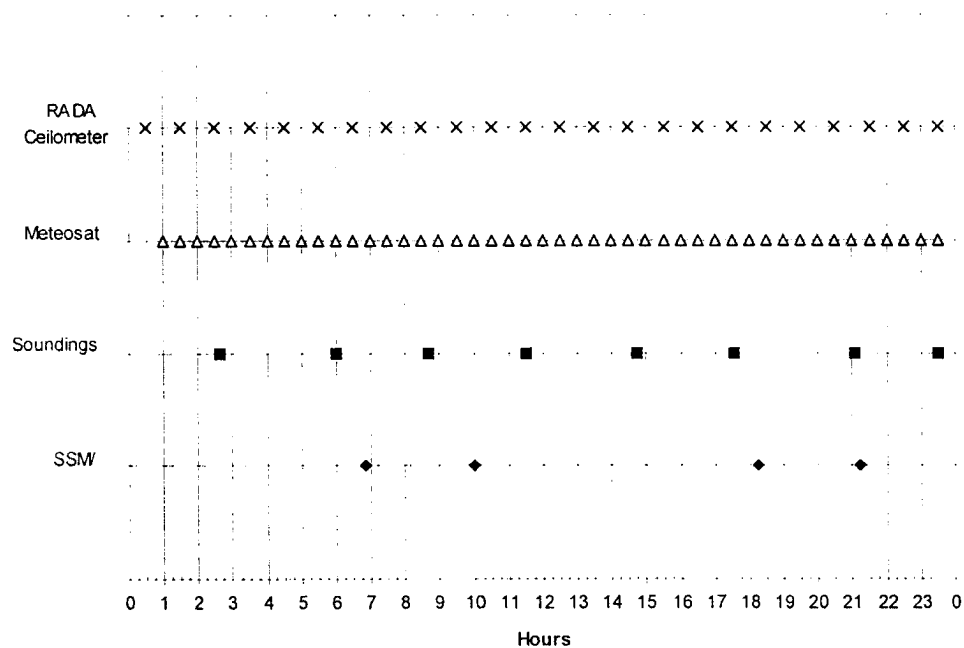


Figure 3.5: Timeline of available data for June 14, 1992 (J116).

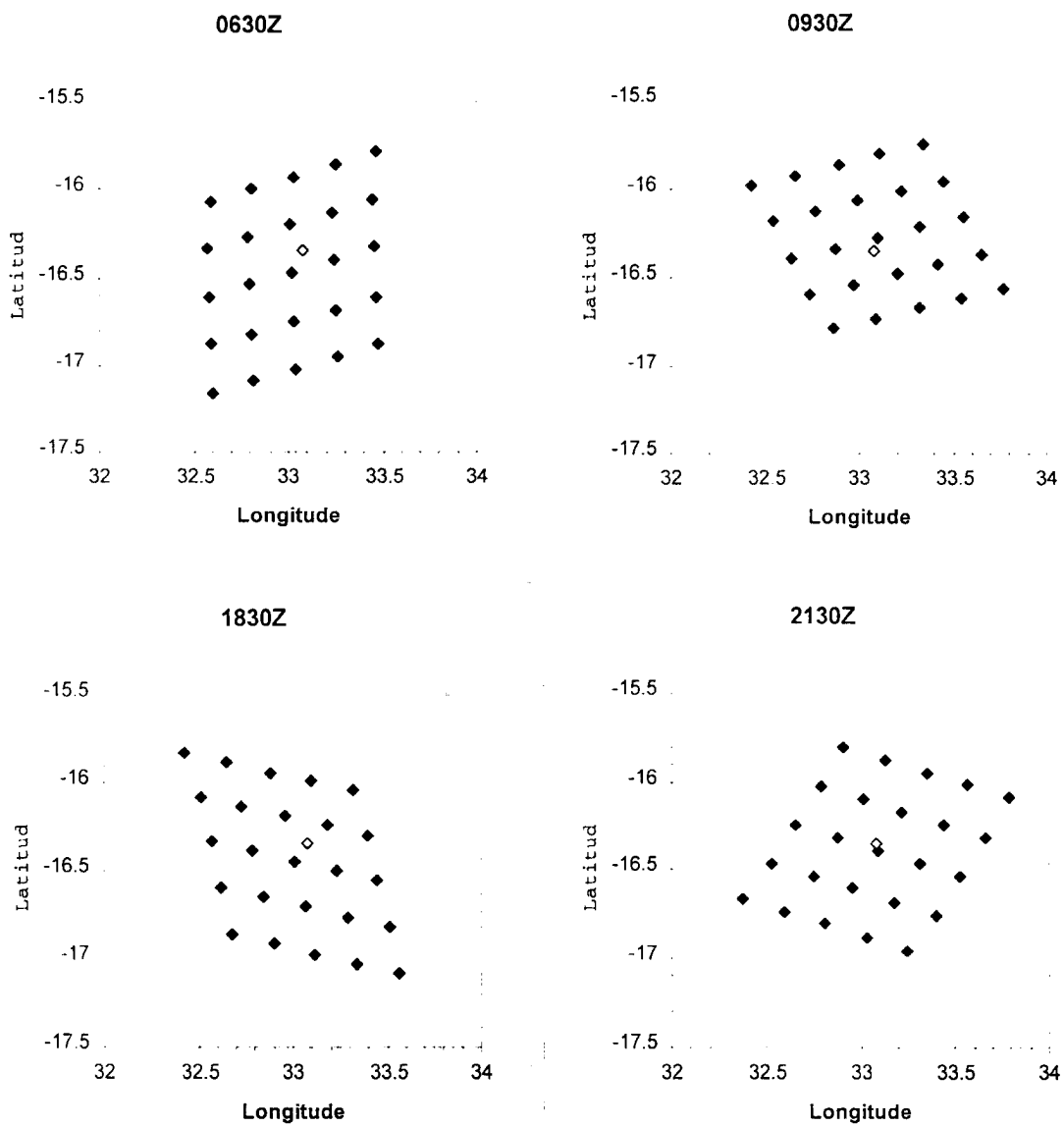


Figure 3.6: Locations of sounding site relative to gridded data points for the different analysis times.

3.2 Satellite data sets

3.2.1 SSM/I

The microwave brightness temperatures used to compute the column integrated liquid water are obtained from the Special Sensor Microwave/Imager (SSM/I) flown on the Defense Meteorological Satellite Project (DMSP) satellites. The satellites are referred to as F-8, F-10, and F-11. Data for this project was obtained from the F-10 and F-11 satellites. The DMSP satellite is in a sun-synchronous orbit at an inclination of 98.8° . The SSM/I instrument performs a conical scan with a 53° viewing angle between the instrument line of sight and nadir. Both vertically and horizontally polarized radiation is measured at 19.35 GHz, 37 GHz, and 85.5 GHz while only vertically polarized radiation is measured at 22.235 GHz. Table 3.1 identifies the resolution of each channel.

Table 3.1: Spatial resolution of the SSM/I channels on the DMSP satellite.

| Channel | Frequency (GHz) | Resolution (km) |
|---------|-----------------|-----------------|
| 1 | 19.35 V | 69 x 43 |
| 2 | 19.35 H | 69 x 43 |
| 3 | 22.235 V | 50 x 40 |
| 4 | 37 V | 37 x 28 |
| 5 | 37 H | 37 x 29 |
| 6 | 85.5 V | 15 x 13 |
| 7 | 85.5 H | 15 x 13 |

Channels 1 through 5 are used to compute the liquid water path as described in section 2.3. Spatial resolution increases with increasing frequency; therefore, the 37 GHz channel

has the smallest resolution of the 5 channels. The 37 GHz channel has 37 km along-track and 29 km cross-track resolution.

The SSM/I brightness temperatures for each channel are obtained from Wentz formatted data. The Wentz format converts the radiation measured by SSM/I into brightness temperature and includes the earth location of each pixel in latitude and longitude coordinates.

3.2.2 GMS

The visible imagery and IR brightness temperatures for Case I are obtained from a Geostationary Meteorological Satellite (GMS). The GMS-4 is a spin-stabilized, Japanese satellite in geostationary orbit located at 140°E. The satellite contains a Visible and Infrared Spin Scan Radiometer (VISSR). The visible channel (0.5 μm - 0.75 μm) has a resolution of 1.25 km at the satellite sub-point and the infrared channel (10.5 μm - 12.5 μm) has a resolution of 5.0 km.

The IR data are used as input into the Greenwald method to determine an average cloud temperature. In addition, the IR and VIS data are used to analyze the cloud scene as described in section 2.1.1. The IR data are also used to indicate cloud top temperature and the height of the highest cloud layer.

3.2.3 Meteosat

Data from the European Space Agency geostationary meteorological satellite, Meteosat-3, was used for Case II. The Meteosat satellites detect radiation in three spectral bands: 0.5 μm - 0.9 μm for the visible band, 5.7 μm - 7.1 μm for the infrared water vapor,

and 10.5 μm - 12.5 μm for the thermal infrared band. Data from the visible and infrared bands were utilized for this research. The visible imagery has a spatial resolution at the sub-satellite point of 2.5 km and the infrared imagery has a resolution of 5 km.

As with the GMS data, the Meteosat infrared imagery data was input into the Greenwald method to determine the cloud liquid water and the visible and IR imagery was used to identify the cloud scene.

3.2.4 Sea surface data

Sea surface temperature (SST) fields were extracted from a weekly averaged, one-degree, gridded, global data set. The SST data set blends skin temperatures retrieved from the Advanced Very High Resolution Radiometer (AVHRR) flown on polar orbiting satellites with SST from ship and buoy observations [Reynolds and Smith, 1994].

3.3 Satellite data set remapping

The microwave, infrared, and sea surface data sets needed to compute cloud liquid water come from different sources and projection spaces. The Polar Orbiter Remapping and Transformation Application Library (PORTAL) [Jones and Vonder Haar, 1992] is used to remap the data to a common projection space prior to computing the cloud liquid water using the Greenwald method.

The concept of PORTAL focuses on processing each data set into a generalized data format (GDF). Once a data set is in GDF format it can be remapped to match the projection space of another data set, also in GDF format, by merging the two GDF files. The GDF file restructures the data and adds a header but retains all the data resolution

and associated earth-location information from the original data set. The header contains information about the type of data and allows the user flexibility to handle conditions like missing data and multiple channels. The data is structured into scan lines to accommodate the time and earth-location of each data point.

For this project, the IR and Sea Surface Temperature (SST) data sets were remapped into the projection space of the SSM/I data set.

CHAPTER 4

CASE I

4.1 Analysis of clouds in scene

4.1.1 Cloud type from satellite VIS and IR imagery

Manual inspection of the visible and IR images for 2345 Z on J356 (figures 4.1 and 4.2, respectively) reveal the cumuliform type clouds that are typical in the tropics. Widespread deep convection south of the rawinsonde sites is related to the ITCZ. Deep convective clouds are identified by very high albedo (white) on the visible image and low temperature (inverted to appear white) on the infrared image. An overlying cirrus layer exists over large portions of the area, appearing white with a wispy texture in both. The storm to the southwest of PSKAV is dissipating, as evident by the cirrus blowoff. Two of the sites, PSN05 and PSMOA, are located near some small dissipating convective storms with a layer of cirrus above. The island site, PSKAV, appears to be located in a more an area of broken cloudiness as evident by warmer, lower level clouds.

4.1.2 Vertical cloud structure over selected sounding sites

Satellite imagery only shows the top layer of clouds. The base of the cloud and any intervening layers are determined here using sounding data from each of the selected sites. The cloud layers are determined from the rawinsonde humidity profile using the threshold method described in section 2.1.2. To supplement these data, the average IR equivalent brightness temperature measured by GMS is computed from a frequency distribution of the 25 gridded data points surrounding each sounding site. The IR

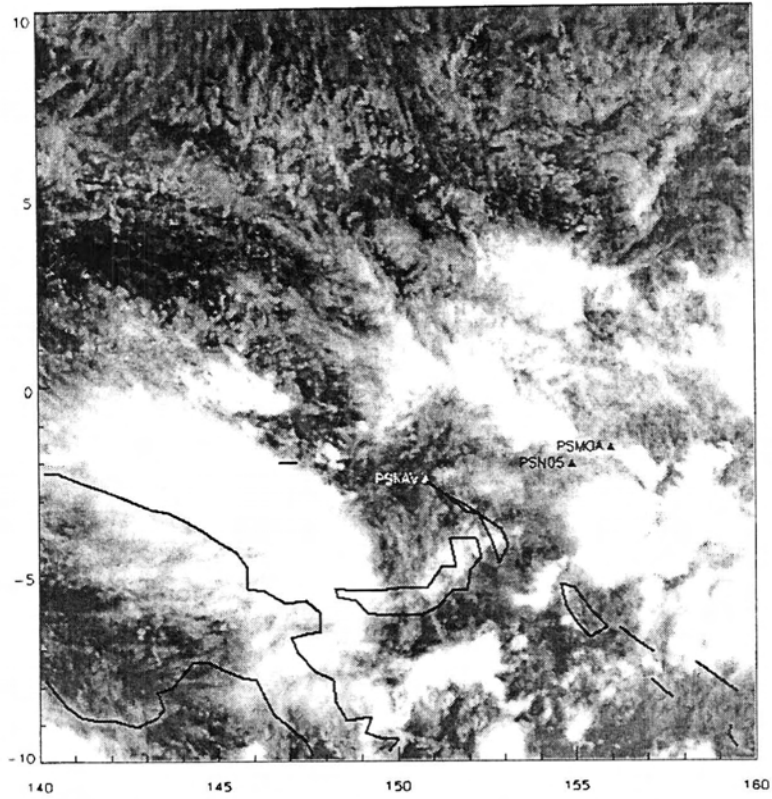


Figure 4.1: GMS visible image for J356 at 2315Z.

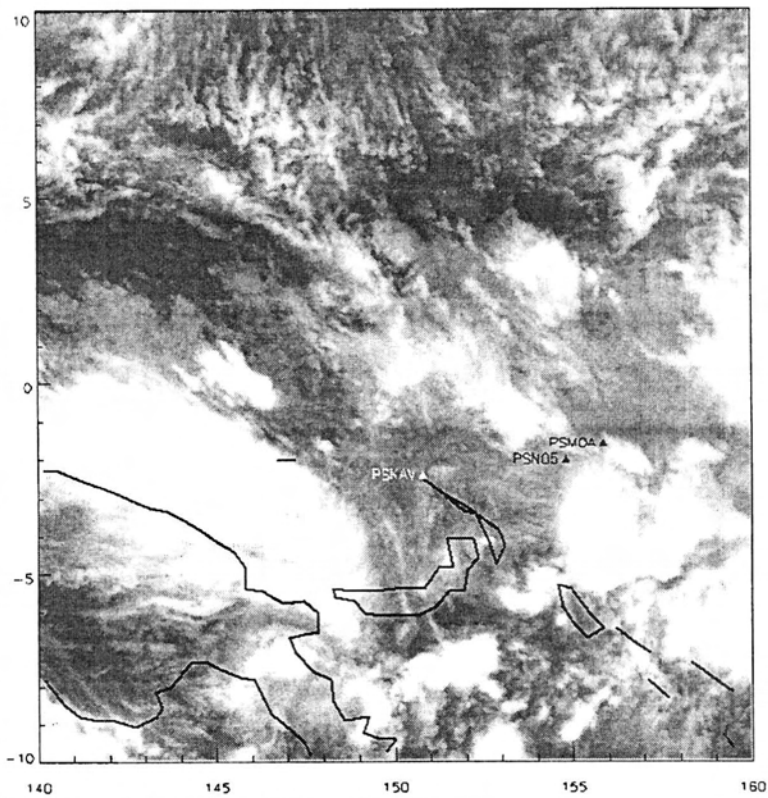


Figure 4.2: GMS infrared image for J356 at 2315Z.

brightness temperature is used in conjunction with the rawinsonde profile to determine the height of the top of the highest cloud layer. This averaged cloud top is used for situations where the RAOBS indicate a cloud top that extends to temperatures lower than $-40\text{ }^{\circ}\text{C}$ since humidity measurements are unreliable at these temperatures.

The vertical humidity profile for each site and time is shown in figure 4.3 with an overlay of the vertical location of the cloud layers.

4.2 Liquid water retrieval

4.2.1 Horizontal distribution of liquid water over scene

Figure 4.4 is a plot of the column integrated cloud liquid water path computed for J356 using the Greenwald method and microwave images from SSM/I and infrared images from GMS. The measured liquid water paths are shown scaled from 0 kg/m^2 to 0.32 kg/m^2 . White is used within the SSM/I swath to correspond to liquid water paths greater than 0.32 kg/m^2 and includes areas where the 37 GHz signal is saturated due to rainfall or the presence of land.

The intensely cloudy areas of the visible and infrared images in figures 4.1 and 4.2 compare well to the areas that contain the highest amounts of liquid water content. The storm regions to the northwest of PSN05 and PSMOA and over the island south of PSKAV (New Britain) have liquid water paths exceeding 0.32 g/m^2 ; however, most of the 37 GHz signal is saturated here due to the presence of rainfall or land. Areas that have less cloud liquid water correspond to the clearer regions of the GMS imagery. There is good correlation between the nearly zero liquid water and the clear areas north and south of PSKAV.

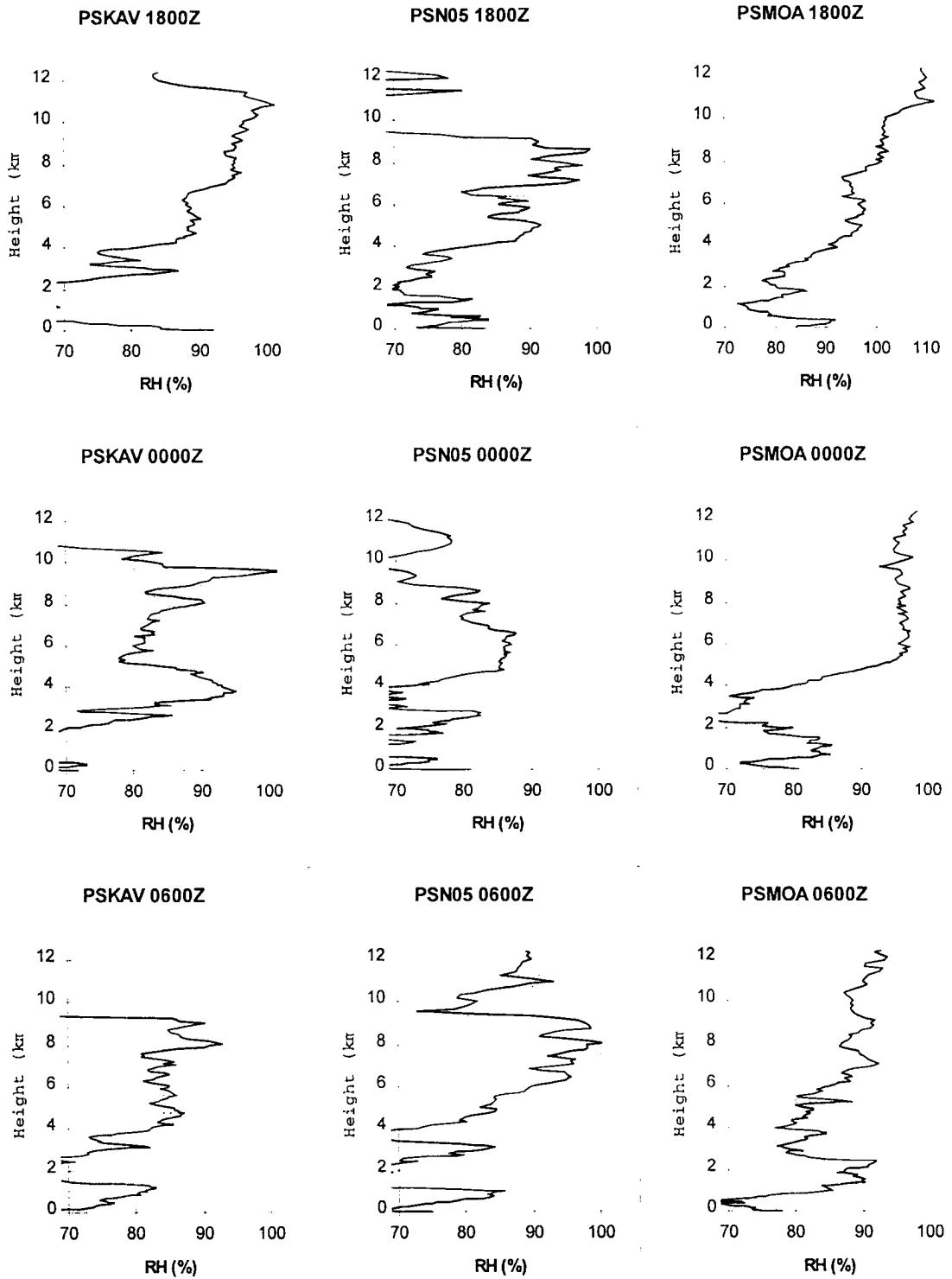


Figure 4.3: Relative humidity vs. pressure overlaid with the vertical location of cloud layers.



Figure 4.4: Liquid water path retrieved from SSM/I for J356.

There is much uncertainty to the retrievals of liquid water in the tropics due to the presence of rain and thick cirrus associated with areas of tropical convection. A reliable measurement of cloud liquid water is not possible when rain is widely distributed within the field-of-view of the SSM/I instrument. Raindrops greatly absorb microwave radiation with the absorption characteristic depending much more on the size distribution of the drops than it does with drops of radius less than 100 μm . Raindrops interact with microwave radiation mostly outside of the Rayleigh regime, so scattering also becomes important. The ice in thick cirrus clouds also impacts the liquid water measurements because ice scatters microwave radiation.

4.2.2 Vertical distribution of liquid water

An average of the 25 pixels surrounding each site is used as the observed columnar amount of liquid water over that site. This water amount is then distributed among the layers of clouds according to the height and type of cloud observed through sounding and satellite data. Cloud height is used to decide whether the cloud contains only water or a mixture of water and ice. A cloud layer with a temperature above 0°C consists of water only. Supercooled water can exist in clouds at temperatures below 0°C but the likelihood of ice increases as the temperature decreases. As seen in figure 4.5, the fraction of clouds that contain ice approaches 100% at -20°C ; therefore, cloud layers that exist between 0°C and -20°C can be considered mixed water/ice clouds and clouds at temperatures below -20°C consist almost entirely of ice. Since ice scatters microwave length radiation, the ice water content is not determined from SSM/I data. The amount of

Figure 4.5: Likelihood of ice with cloud temperature [from Rogers and Yau, 1994]

ice in the clouds is calculated here using an empirical formula that relates ice water content to temperature [Liou, 1986]

$$\ln(IWC) = -7.6 + 4 \exp \left[-0.2443 \times 10^{-3} (|T| - 20)^{2.445} \right] \quad (4.1)$$

The type of cloud gives an indication of the approximate liquid water content that it should have. Many observational studies have used aircraft to measure the droplet size distribution in various types of stratiform and cumuliform clouds [Paltridge, 1974 and Slingo *et al.*, 1982, for example]. The size and number concentration of drops determines the liquid water content, or mass of cloud water per unit volume, according to

$$LWC = \frac{4}{3} \pi \rho_l \int_{\Delta r} r^3 n(r) dr \quad (4.2)$$

where ρ_l is the density of water. A summary of the drop size distributions determined by various investigators is shown in table 4.1.

Table 4.1: Drop size distribution and LWC for different clouds [from Liou, 1992]

A method for distributing water among the cloud layers is now described. Clouds are first categorized as water clouds, mixed water/ice, or ice clouds by temperature only. A cloud that covers two categories is split into those two types of cloud. For example, a cloud with a base temperature above 0 °C and a top temperature below 0 °C is split into a water cloud up to the 0°C level and a mixed water/ice cloud above that level. Next, the liquid water profile is computed using the height of the cloud and an initial, reasonable value for the liquid water content. Liquid water path is integrated liquid water content

$$W = \int_b^{\Delta z} w dz \quad (4.3)$$

where w is the liquid water content in g/m^3 and Δz is the thickness of the cloud in meters. The liquid water content is adjusted until the total liquid water path from each cloud equals the total liquid water path. Figure 4.6 shows the distribution of liquid water that is

used for the clouds at each site and analysis time. These cloud heights and liquid/ice water contents are key inputs to the heating rate calculations as shown earlier in section 2.2.5.

One condition that was shown in the sensitivity study but not considered here is that liquid water is highly variable within a cloud. In a cumulus cloud, the amount of liquid water peaks in the upper half a cloud, with smaller amounts at the base and top of the cloud [Pruppacher and Klett, 1978; Starr and Cox, 1985; and Spinhirne *et al.*, 1989]. Although this vertical profile of liquid water within a cloud is important to accurately compute heating and cooling rates [Stephens, 1978], a constant-height distribution of liquid water is used here because this investigation looks at the change in heating rates integrated over 200 mb layers.

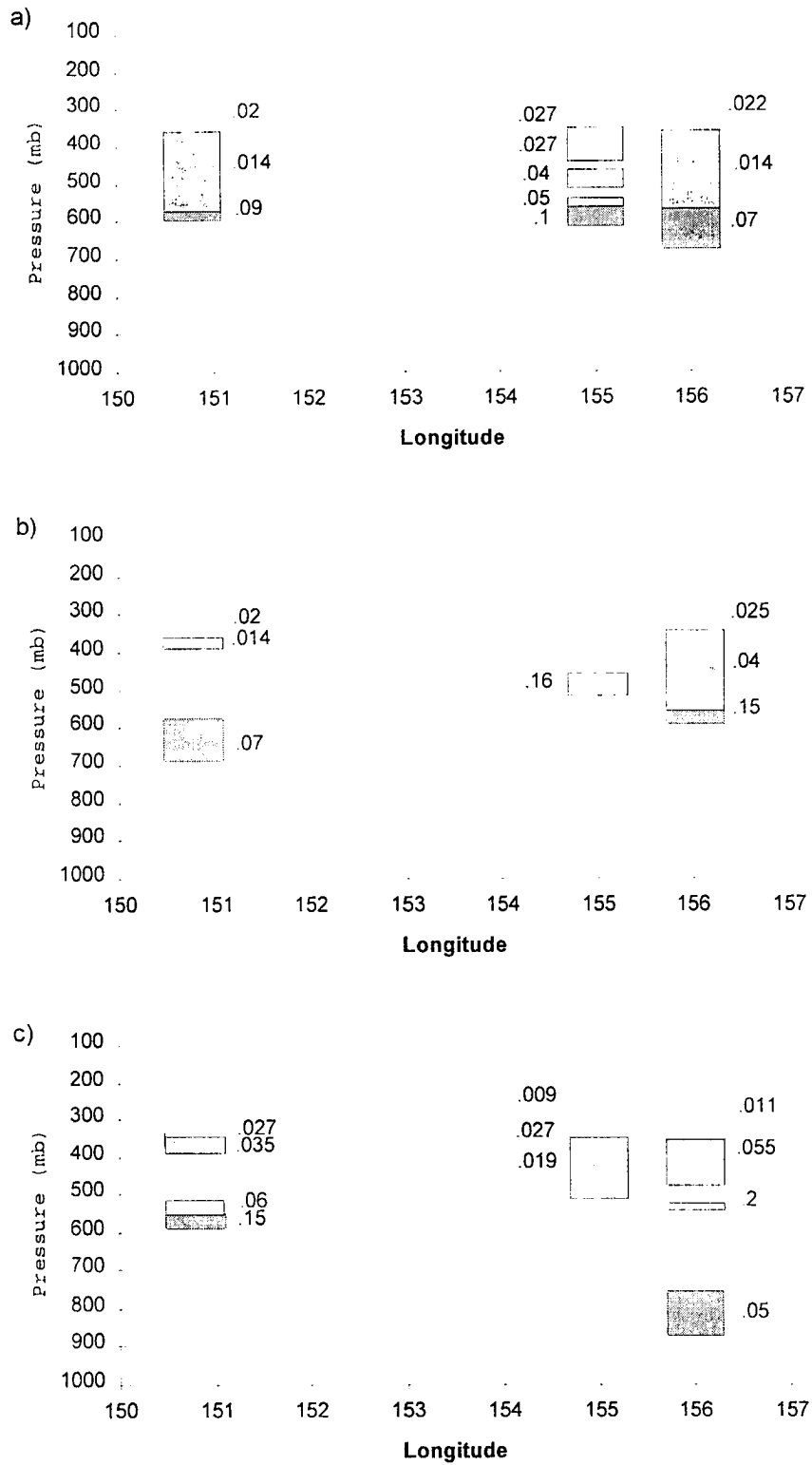


Figure 4.6: Distribution of LWC for cloud layers: a) 1800Z, b) 0000Z, c) 0600Z.

4.3 Heating and cooling rates

4.3.1 Analysis

The LWBAND program is used to calculate the infrared heating and cooling rates for each site and analysis time. This program computes the net flux and heating rate with a vertical resolution of 10 mb. The net flux is dependent on cloud optical depth which is a function of cloud thickness and the LWC or IWC of the cloud. There was only one time (0000Z) that all three sites were in the field-of-view of the SSM/I during the selected analysis time period; therefore, only data from this time is available to compute the LWP.

The vertical profiles of the heating and cooling rates for each site and analysis time are shown in figures 4.7 through 4.9. Sharp peaks in heating or cooling rate occur at the cloud base and top, respectively, for each layer. The largest cooling rate ($-100\text{ }^{\circ}\text{K/day}$) and heating rate ($35\text{ }^{\circ}\text{K/day}$) occur for the single-layer, mixed water/ice cloud over PSN05 at 0000Z. The heating and cooling profiles for the multiple-layer and divided single-layer clouds have maximum heating and cooling rates that nearly balance each other. The maximum cooling rates for these clouds range from -20 TO $-30\text{ }^{\circ}\text{K/day}$, and the heating rates range from 15 to $30\text{ }^{\circ}\text{K/day}$.

In most of the multiple-layer cloud cases, the peak cooling rate for the lower cloud layers is less than the peak cooling rate of the top layer. This suggests that the upper cloud increases the downwelling infrared flux that is incident on the lower clouds, which decreases the net flux and reduces the cooling rate of the lower clouds.

The abrupt changes in heating rate within the divided single-layer cloud layers result from the division of the cloud into the distinct water, mixed, or ice regions. A more

appropriate scheme would be to gradually change the water content with height in the cloud according to the microphysics of the cloud.

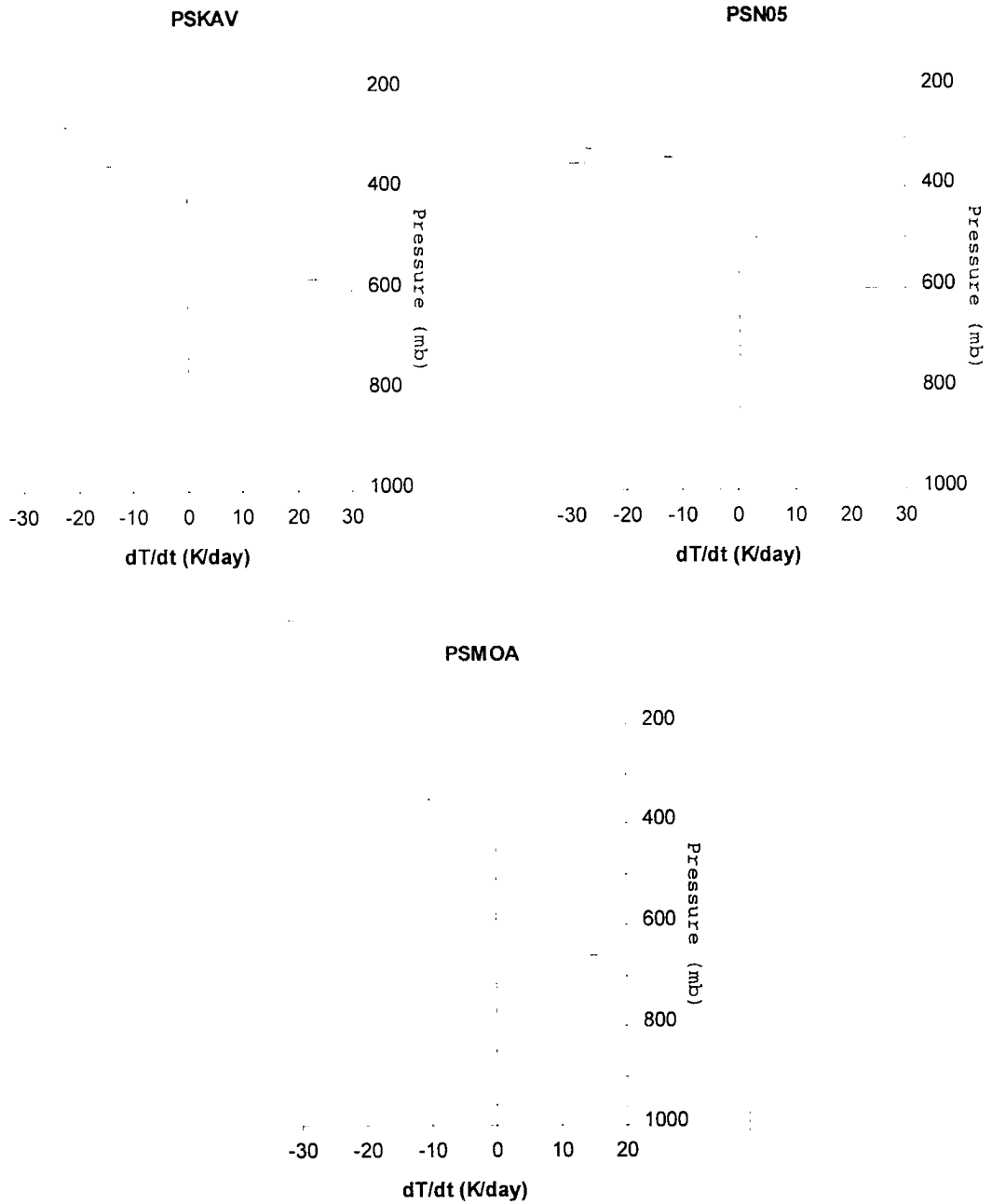


Figure 4.7: Heating and cooling rate profile for each site at 1800Z.

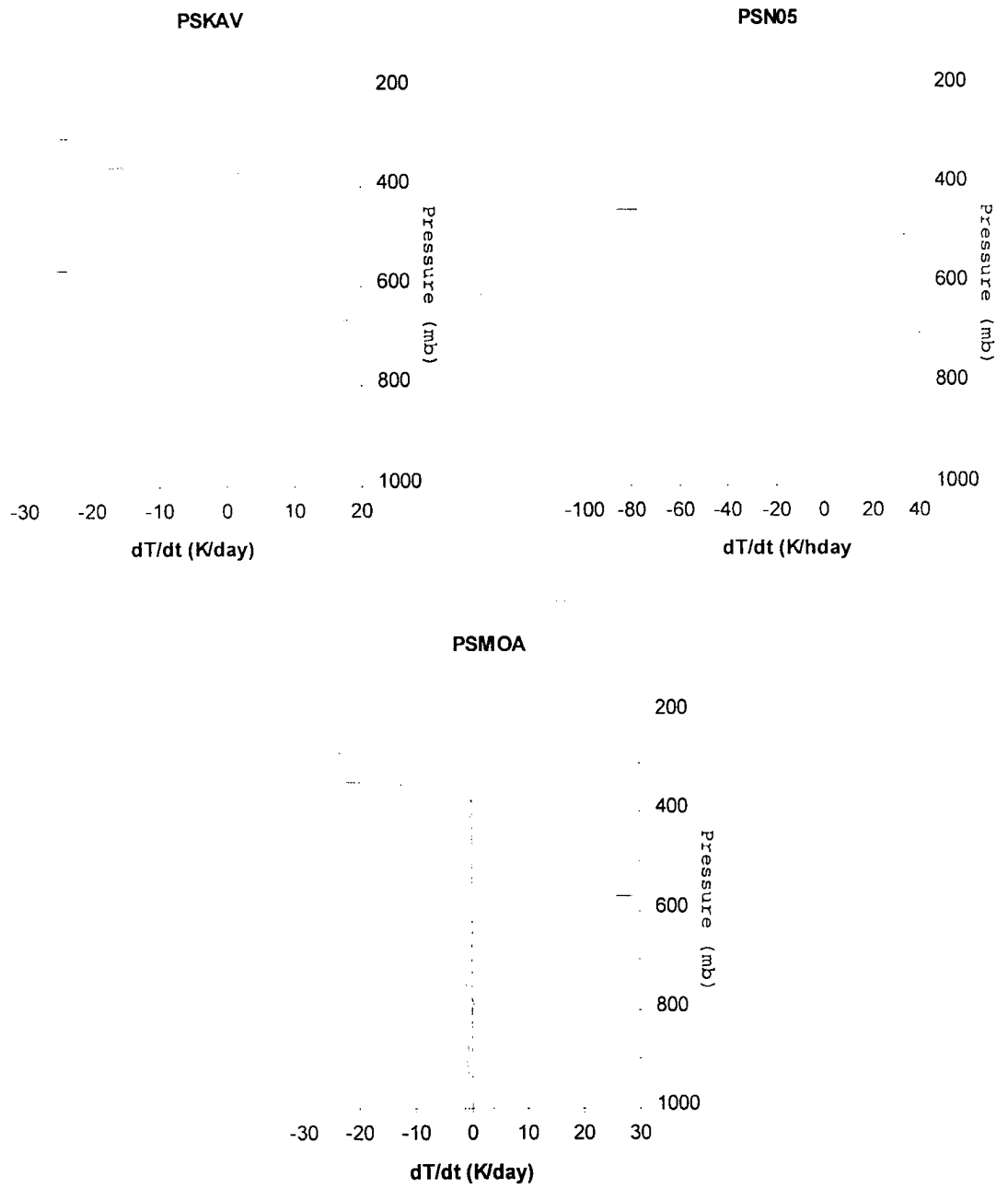


Figure 4.8: Heating and cooling rate profile for each site at 0000Z.

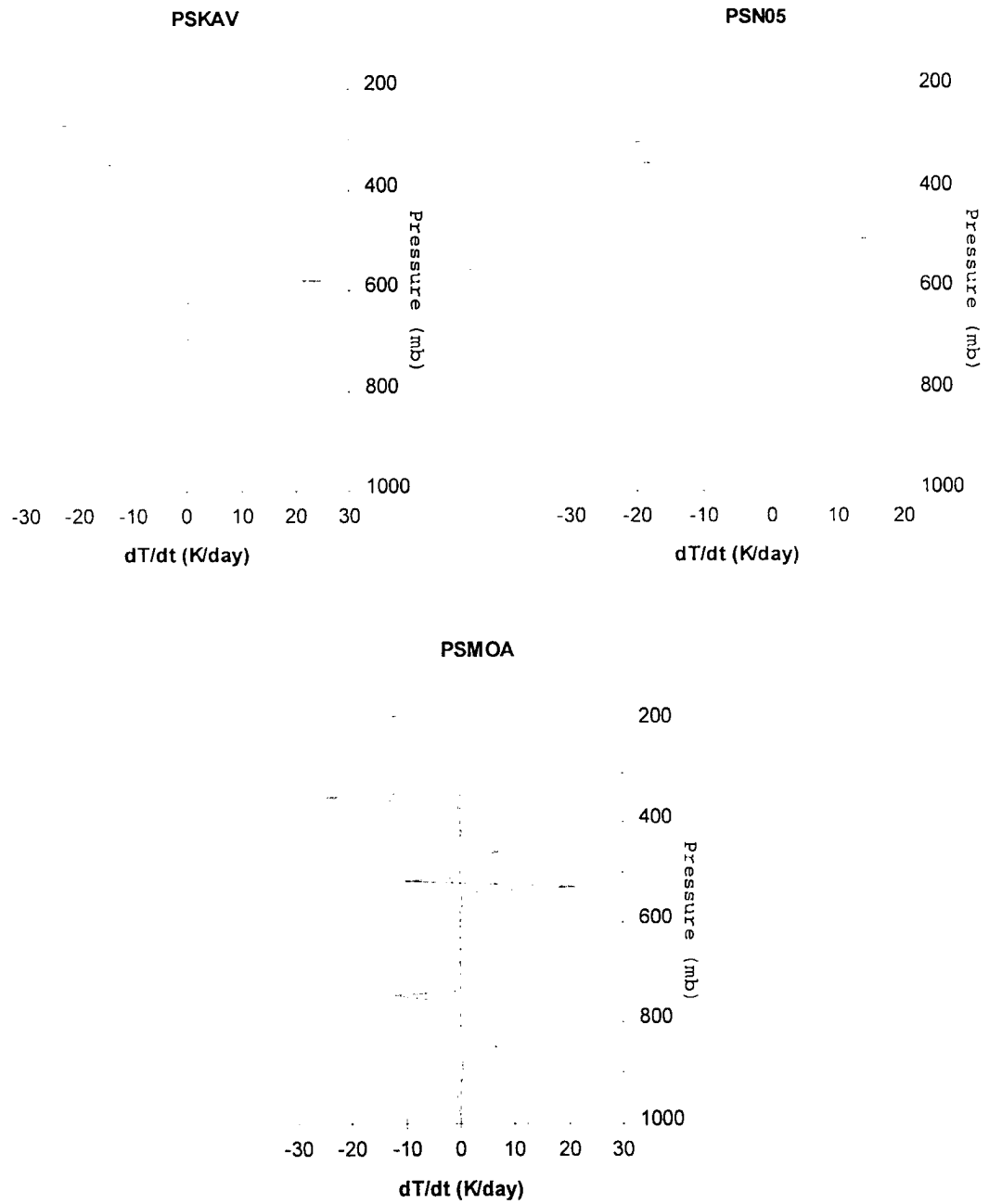


Figure 4.9: Heating and cooling rate profile for each site at 0600Z.

The local, abrupt changes in the heating profile are averaged out when the heating and cooling rates are integrated over a 200 mb atmospheric layer. Figures 4.10 through 4.12 show the integrated heating and cooling rates for each site and analysis time. These plots reveal the vertical atmospheric heating or cooling for the observed cloud scene. Most noticeable are the vertical gradients between the 200-400 mb layer and the 400-600 mb layer for every analysis time which are due to the heating differential between the cloud top and base. On average the cloud top in all the cases is located in the 200-400 mb layer and the cloud base is located in the layer below. The thermal difference between the two layers causes turbulent motion which destabilizes the cloud.

The effect of radiative heating on overall diabatic heating is considered when based on the magnitudes of latent and sensible heating, radiative heating and cooling variations are at least ± 0.2 °K/day/200 mb [WCRP-86, 1994]. Looking at the three analysis times, the highest layer of the atmosphere (200 - 400 mb) experiences cooling rates that range from 0.26 °K /day to 0.63 °K /day. This is not surprising since most of the cloud tops are located here. The 400 - 600 mb and 600 - 800 mb layers have net heating or net cooling, due to the variety of cloud conditions that exist in this region. The magnitude of heating/cooling is small though, less than ± 0.2 °K/day everywhere except PSN05 at 0000 Z where the cloud top is below 400 mb. This situation also results in a horizontal heating gradient, from PSN05 to PSMOA, that could have an impact on the atmospheric dynamics in this area. The lowest layer has a small amount of cooling at all sites except PSMOA at 0600 Z where the base of the low cloud here provides radiative heating of the layer.

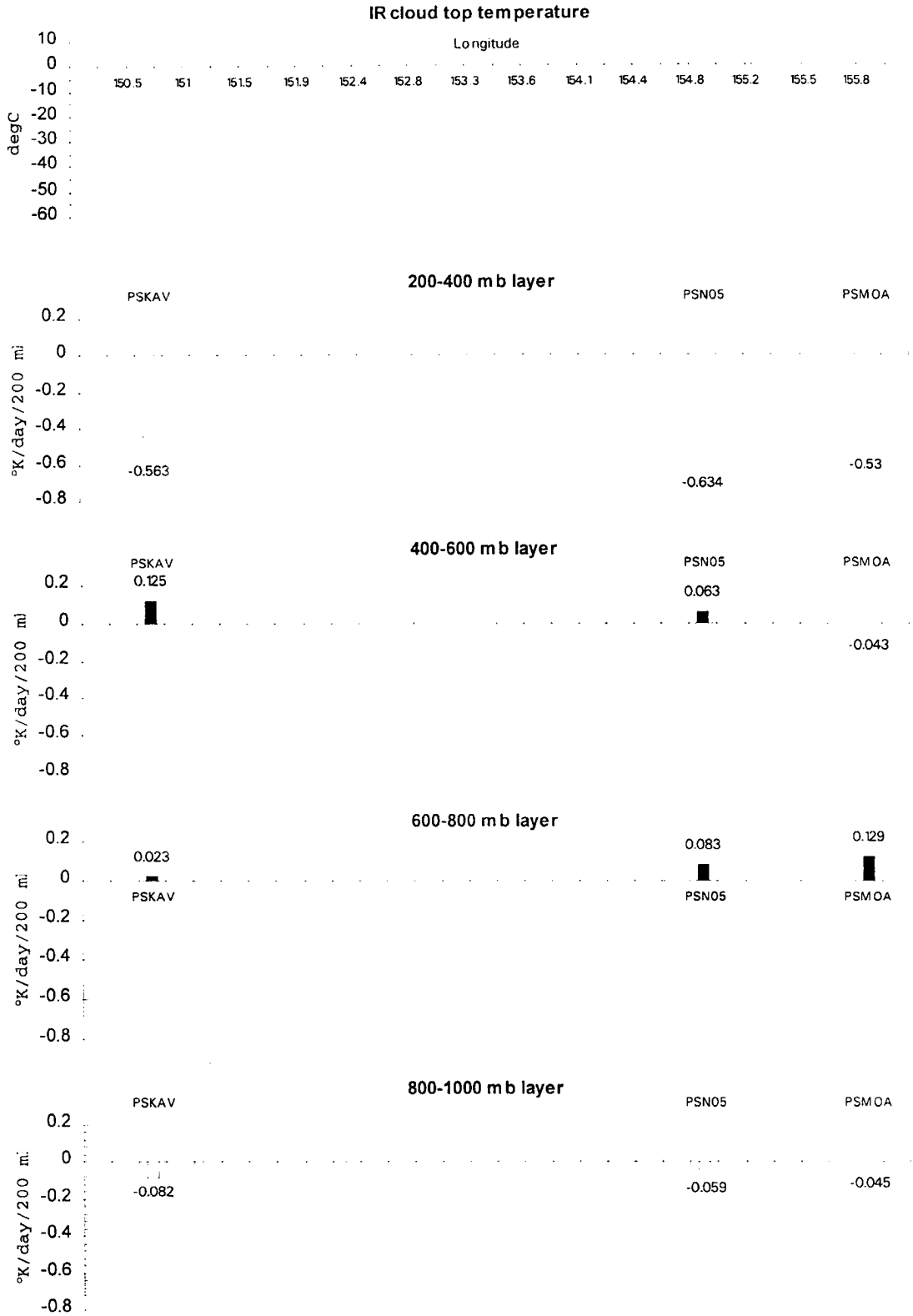


Figure 4.10: Heating rate integrated over 200 mb layers for 1800Z

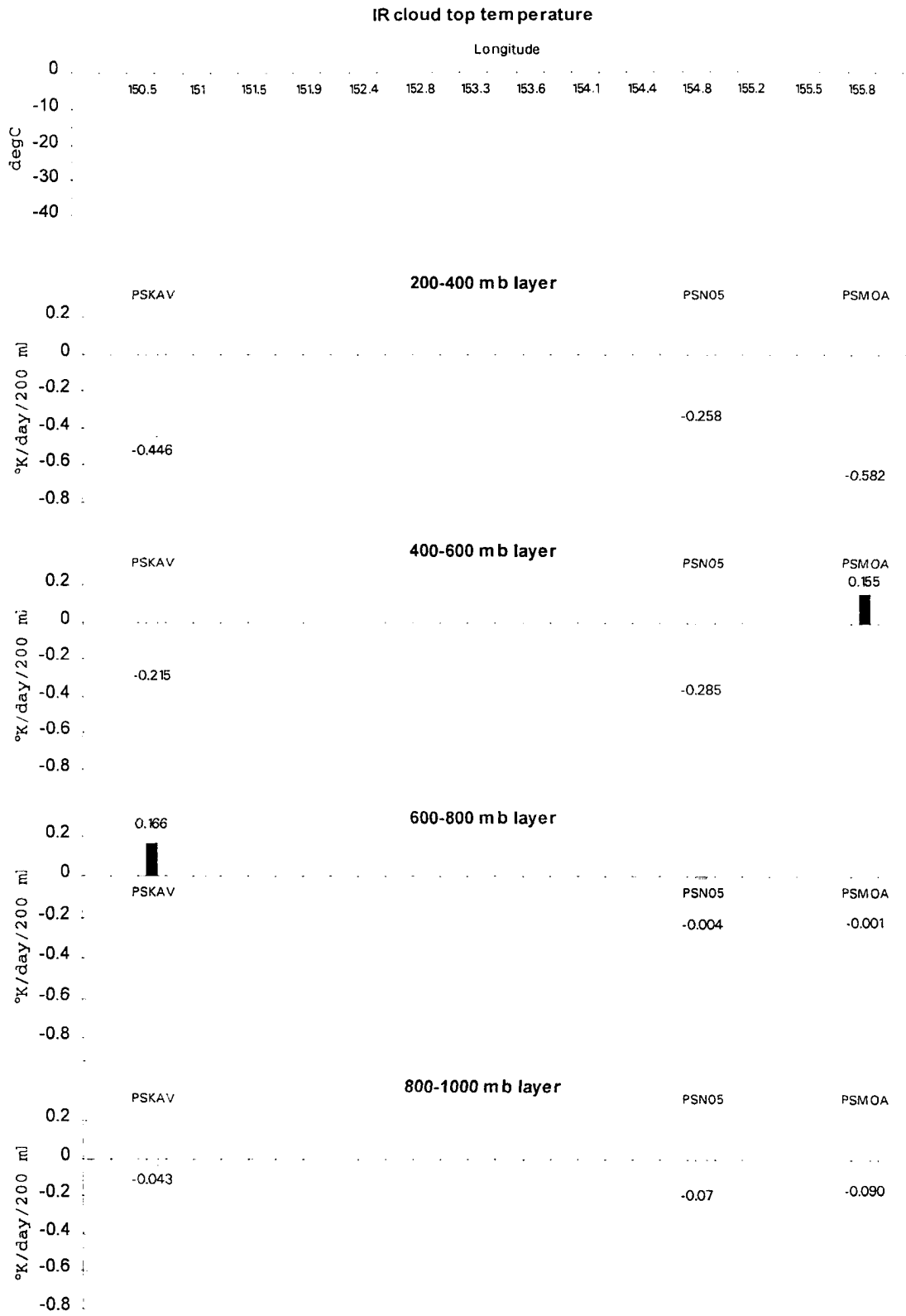


Figure 4.11: Heating rate integrated over 200 mb layers for 0000Z.

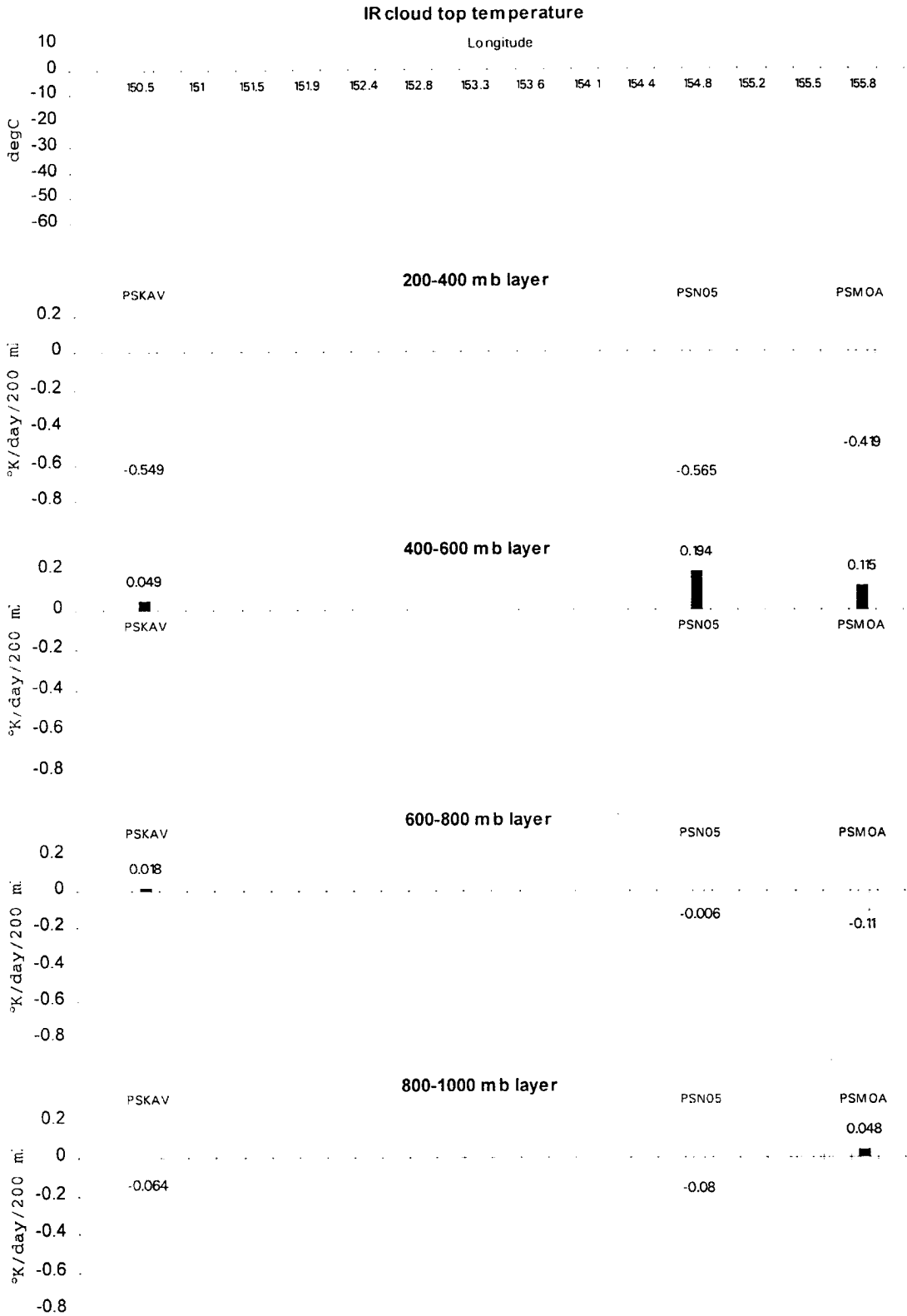


Figure 4.12: Heating rate integrated over 200 mb layers for 0600Z.

4.3.2 Cloud forcing

Next, the overall impact of clouds on radiative heating and cooling rates is investigated by looking at the cloud forcing. The LWBAND radiative transfer program was run with sounding data from 0000Z but without clouds. The resulting heating and cooling profiles are shown with and without clouds in figure 4.13. The profiles show that without the clouds there is cooling throughout the atmosphere and with the cloud, there is warming beneath the cloud and increased cooling above the cloud. The cloudy minus clear heating rates, integrated over 200 mb layers, are plotted in figure 4.14. Negative numbers indicate that the atmosphere is further cooled by the clouds and positive numbers show where the atmosphere is warmed. Figure 4.15 shows the percentage difference. The largest changes occur in the middle layers.

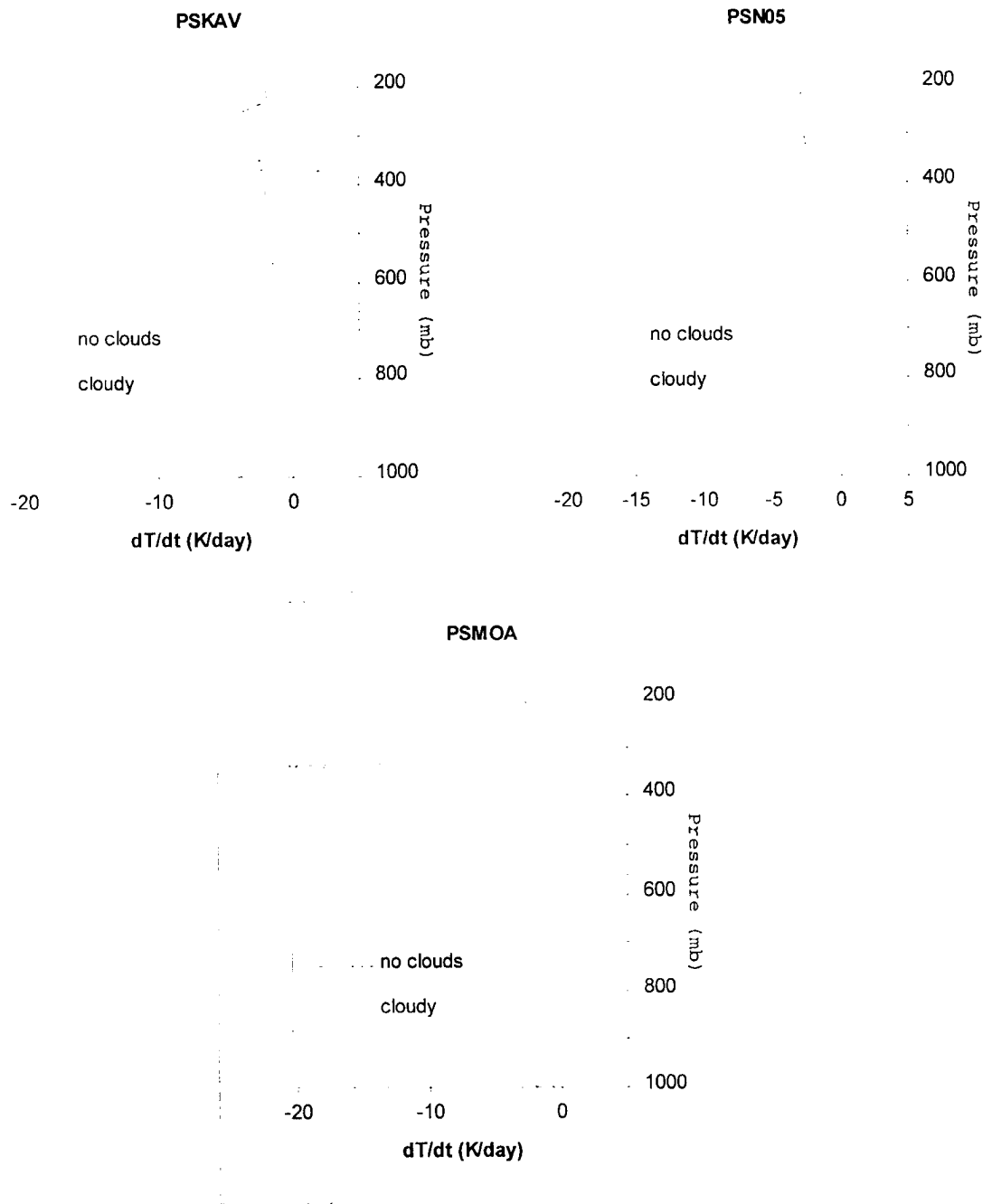


Figure 4.13: Heating and cooling rate profiles with and without clouds at each site at 0000Z..

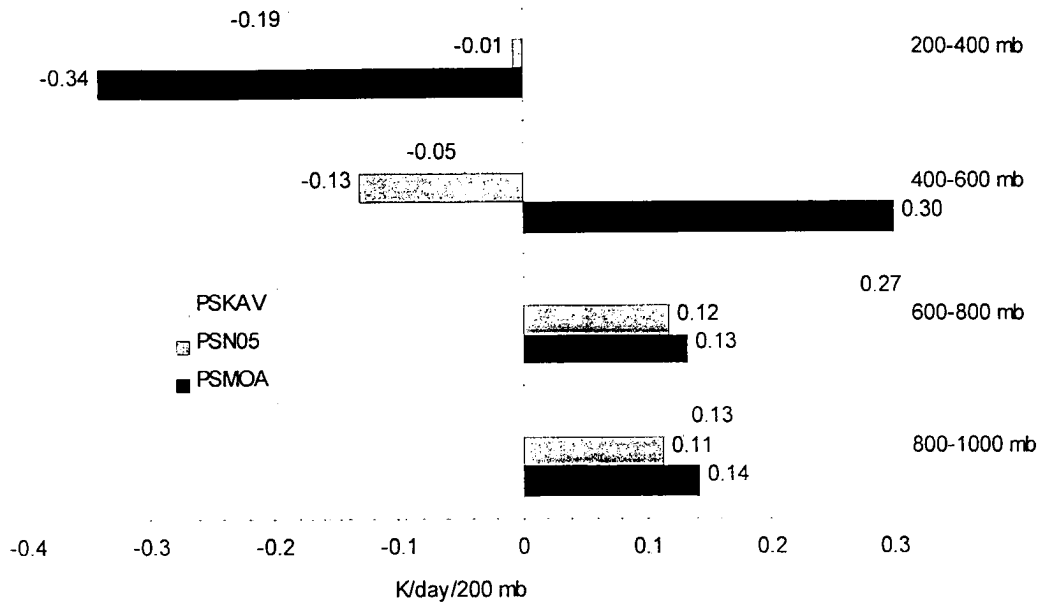


Figure 4.14: Change in heating rate due to cloud (cloudy-clear) for each layer.

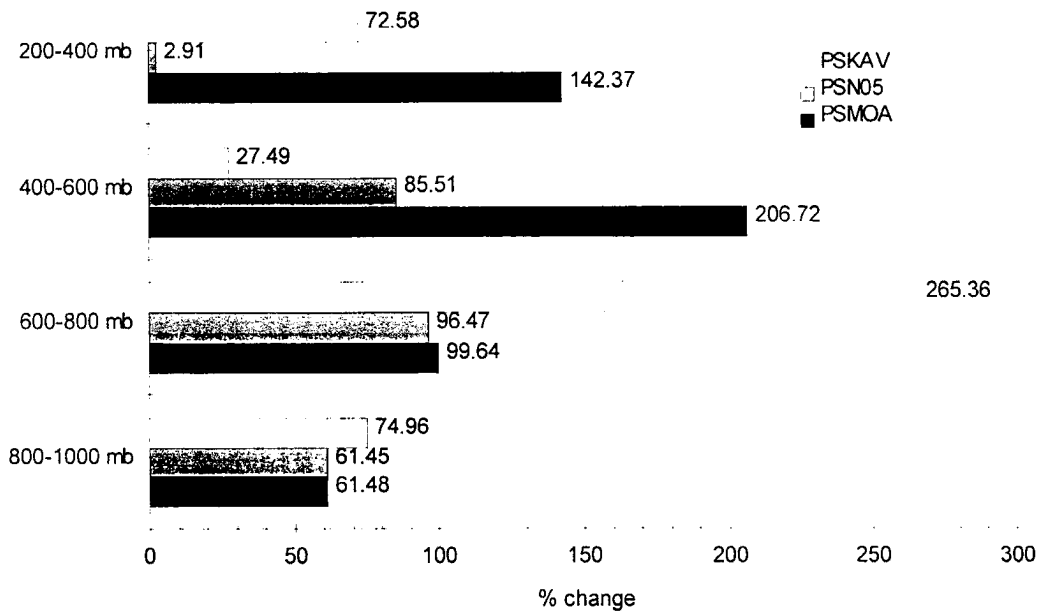


Figure 4.15: Percent change in heating rate due to cloud for each layer..

CHAPTER 5

CASE II

5.1 Analysis of clouds

5.1.1 Cloud type from satellite visible and IR analyses

The visible (figure 5.1) and infrared (figure 5.2) Meteosat-3 images for 0930Z on June 14 show a low, broken stratus cloud deck over the island of Port Santo. The infrared image clearly shows the warmer, lower cloud layer while the visible image shows the stratus-stratocumulus cloud cover over the area. Figures 5.3 and 5.4 show the visible and infrared images, respectively, for 1830Z of the same day. The infrared image continues to show warmer, lower cloud cover. The visible image shows that the stratus clouds are beginning to dissipate.

5.1.2 Vertical cloud structure over sounding site

The vertical profile of the cloud layer is obtained from the 8.7 mm Doppler radar used at ASTEX. Heidinger [1994] used the radar data to compute the bases and tops of the cloud layer on June 14 which he then used in his study of the finite effects of cloud layers on infrared radiation. Figure 5.5 shows that the height and depth of the cloud layer for this day is a maximum in the morning hours, decreases throughout the day, and rises slightly during the nighttime hours.

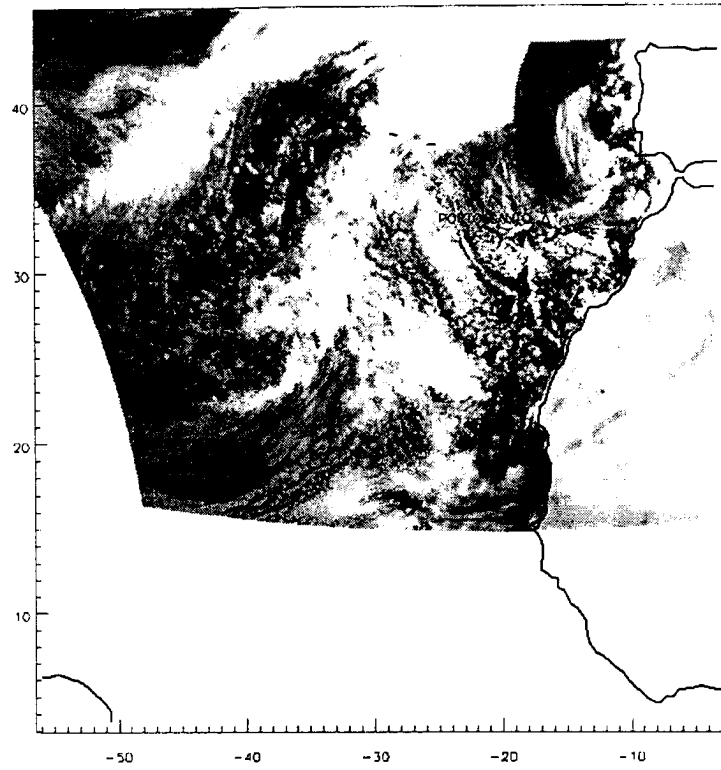


Figure 5.1: Visible image from Meteosat-3 for 0930Z.

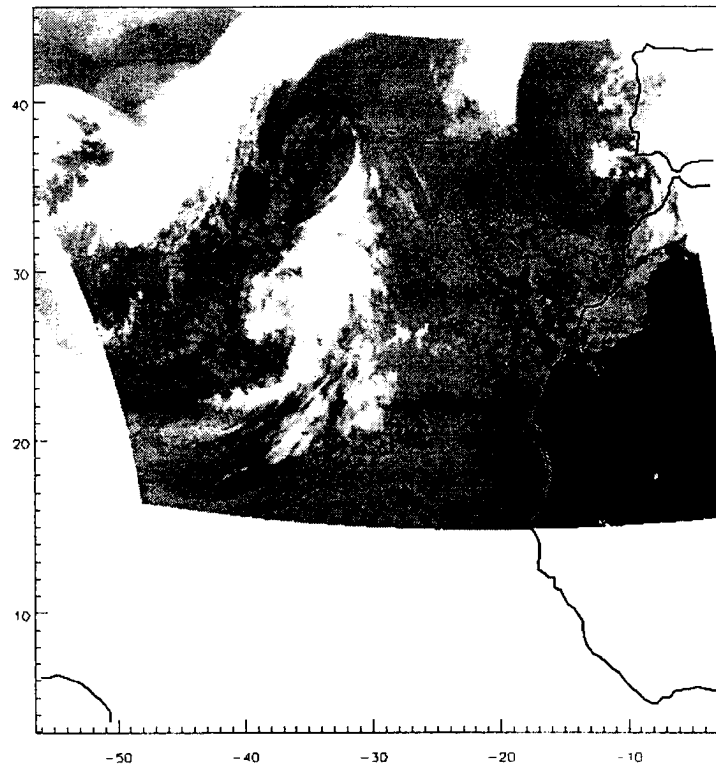


Figure 5.2: Infrared image from Meteosat-3 for 0930Z.

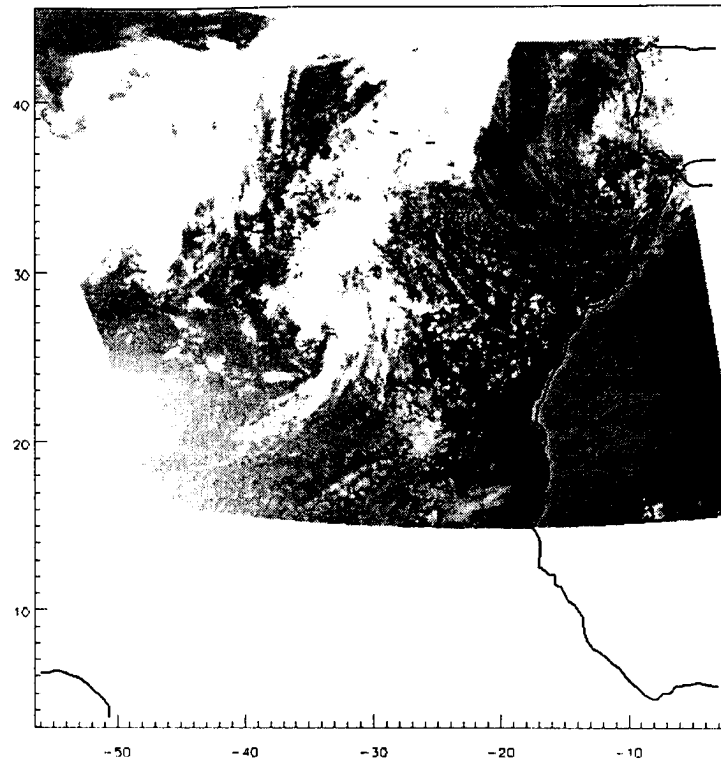


Figure 5.3: Visible image from Meteosat-3 for 1830Z.

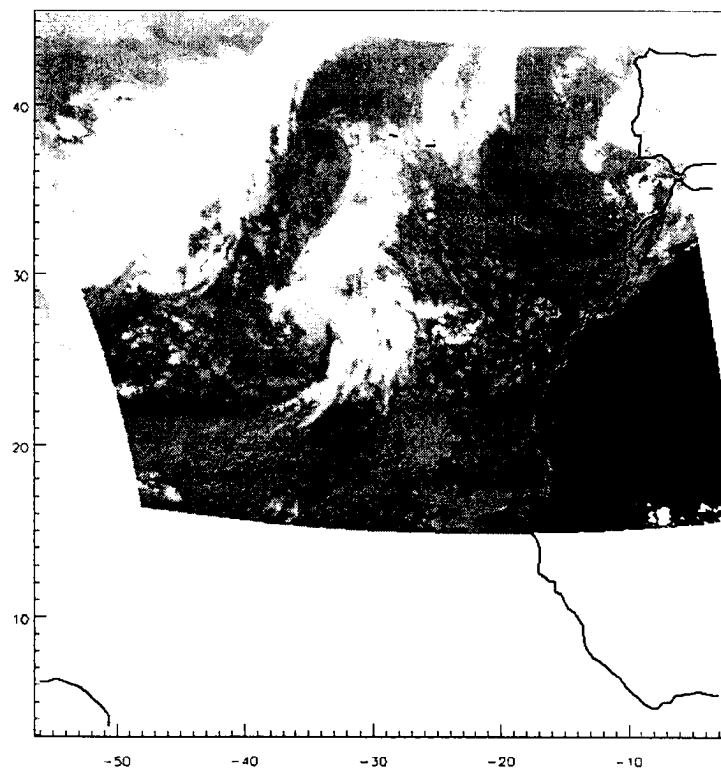


Figure 5.4: Infrared image from Meteosat-3 for 1830Z.

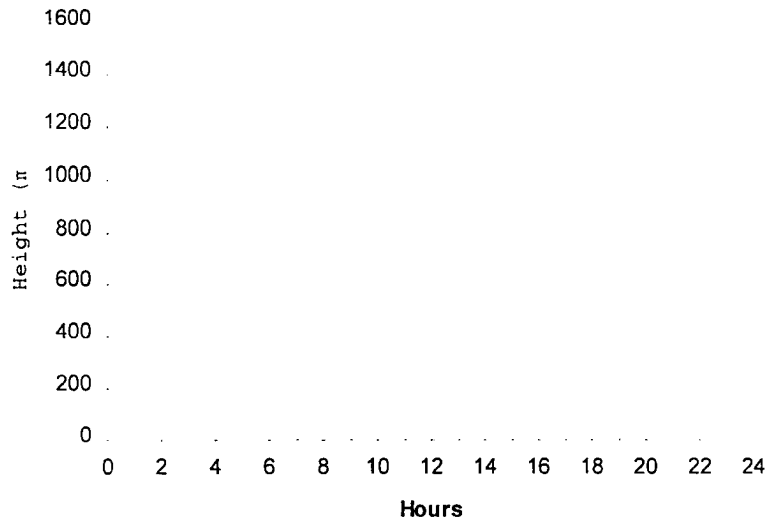


Figure 5.5: Height and depth of cloud layer from radar data, June 14, 1992.

A comparison between the radar derived cloud profile and other methods to determine cloud location is provided here with results shown in figure 5.6. The cloud base and top are computed from the rawinsonde humidity profile data using the threshold method used in case I. The height of the lowest cloud base is determined from hourly averaged ASTEX ceilometer data. The top of the cloud is located from averaged infrared brightness temperatures measured by the Meteosat-3 satellite and remapped using PORTAL to the projection and resolution of the 37 GHz SSM/I channel. The bases and tops from all of these methods are compared to the radar derived cloud height and depth in figure 5.6.

The ceilometer and radar data agree rather well except for the time period from 1000Z to 1600Z when the ceilometer indicates either higher level clouds or a break in the clouds and the radar shows both cloud base and cloud top at about the same height. The

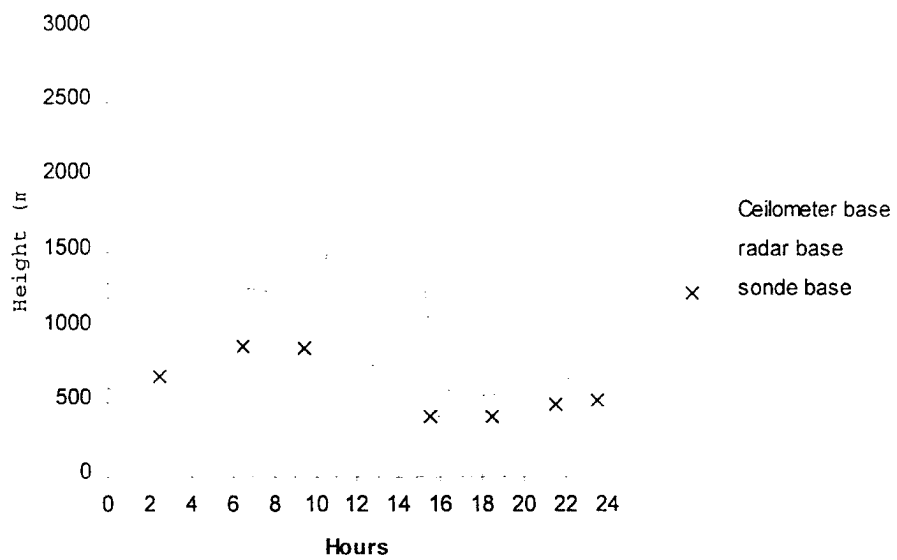
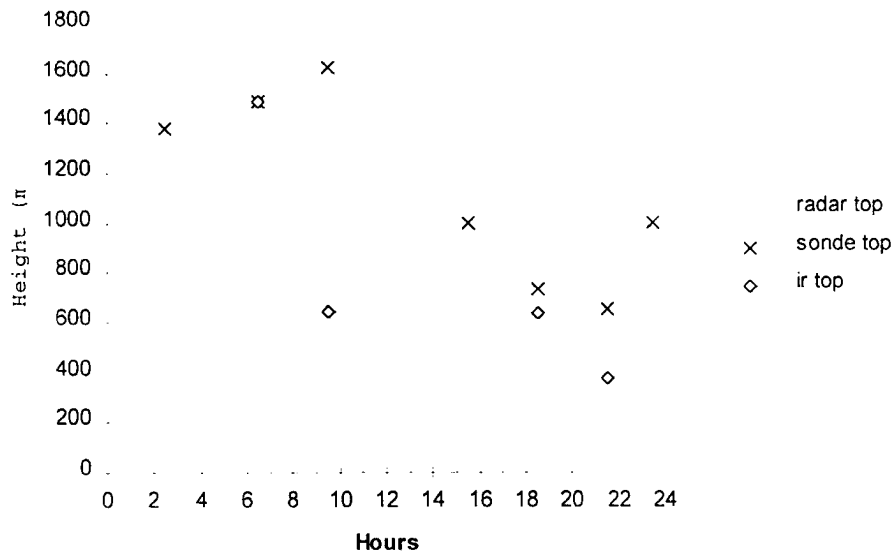


Figure 5.6: Comparison of retrieval of cloud base and top using different methods.

relative humidity from the 1130Z sounding also indicate a lack of cloud cover at this time.

The cloud base derived from the soundings are consistently lower than the radar and ceilometer derived base but follow the trend well. Comparison between the cloud top from sounding data and from radar is slightly better. The top of the cloud is very evident in the soundings (figures 5.7 - 5.10) due to the sharp inversion that occurs above the stratus layer. The cloud top derived from the infrared brightness temperature matches well at 0630Z and 1830Z but is lower at 0930Z and 2130Z.

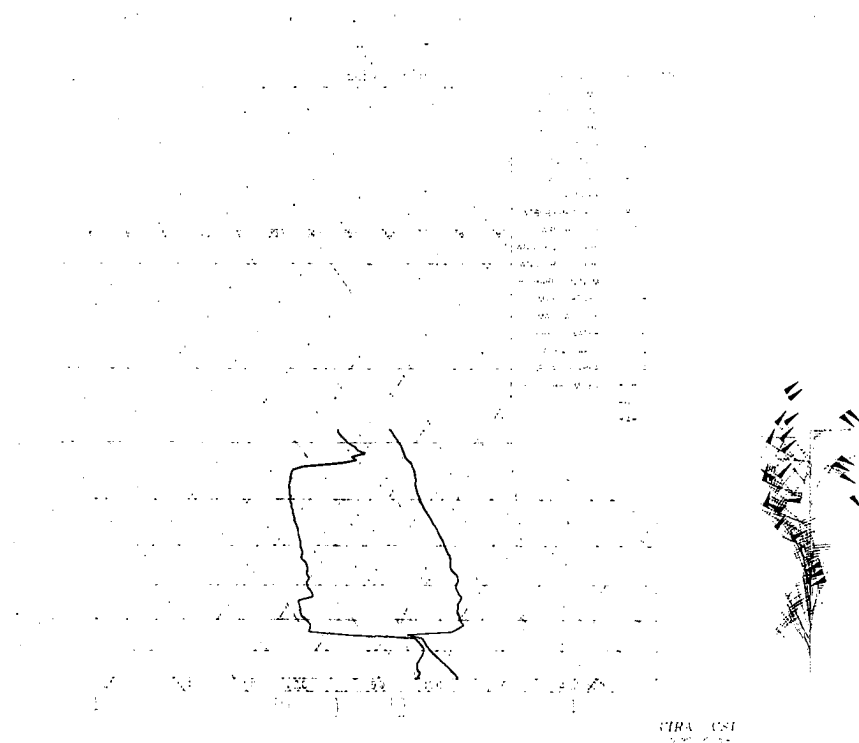


Figure 5.7: Sounding for J166/0638Z.

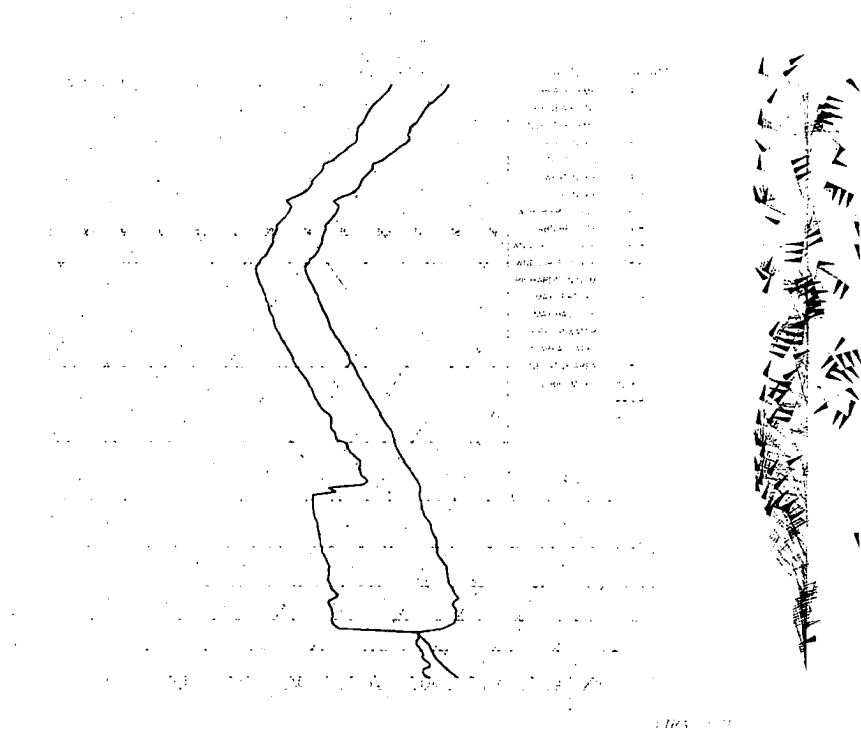


Figure 5.8: Sounding for J166/0824Z.

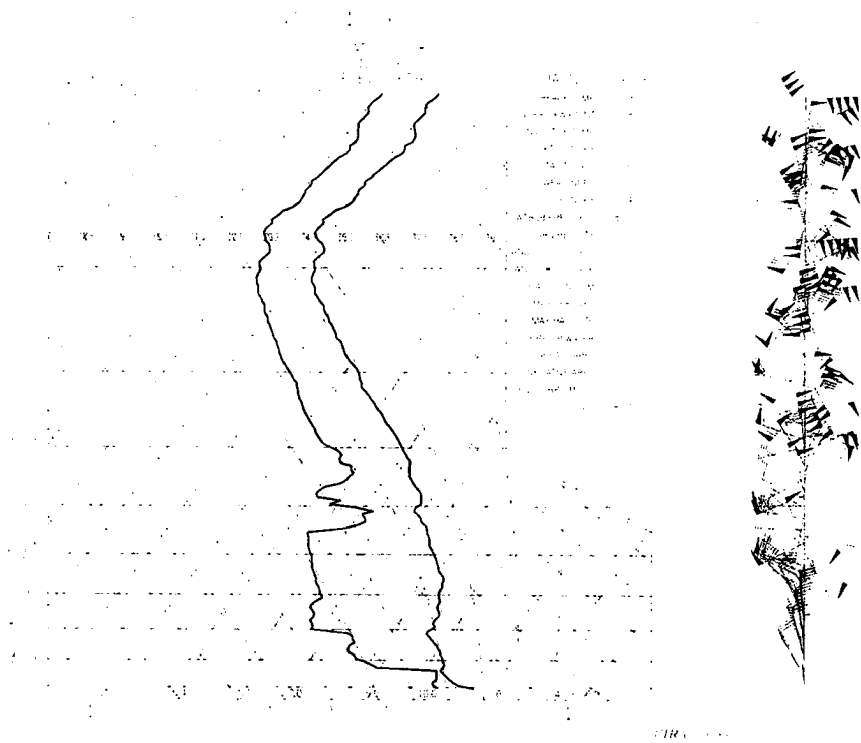


Figure 5.9: Sounding for J166/1735Z.

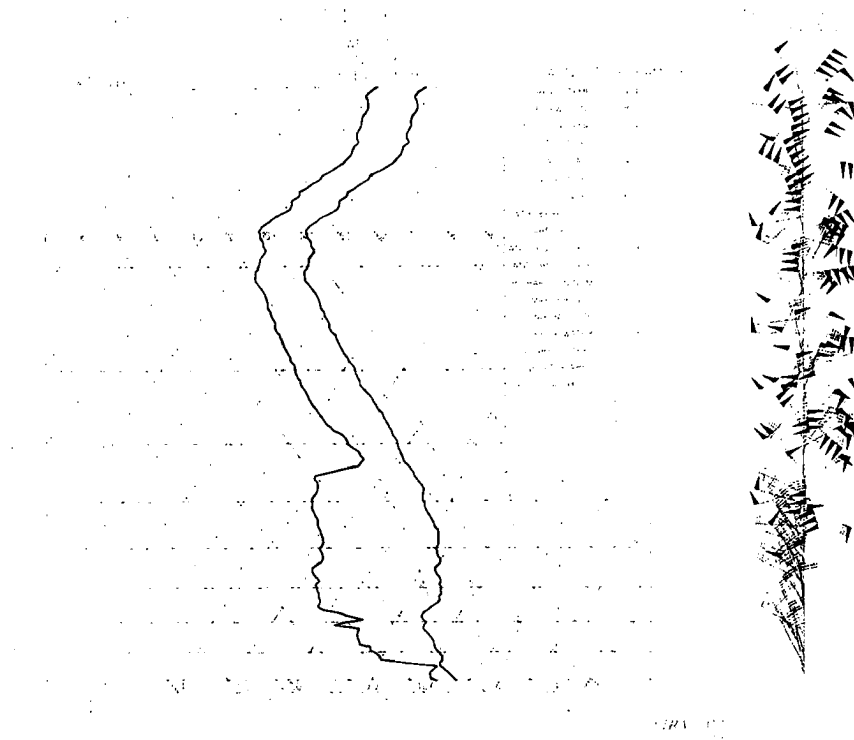


Figure 5.10: Sounding for J166/2000Z.

5.2 Liquid water retrieval

5.2.1 Horizontal distribution of liquid water over scene

The integrated column liquid water paths for the area surrounding Porto Santo are shown in figures 5.11 through 5.14. The liquid water path is computed in the same manner as case I, except infrared data from Meteosat-3 is used. The calculated liquid water path is shown scaled from 0 to 0.16 kg/m^2 in all of the figures with white used to denote liquid water paths greater than 0.16 kg/m^2 for areas within the SSM/I swath.

The retrieved liquid water path over the entire area is low ($< .1 \text{ kg/m}^2$). The “bull’s eye” patch of very high liquid water content that appears on all of the cloud liquid water plots is most likely due to errors in the microwave retrieval over the island of Porto Santo. There is some correlation between increased cloudiness and high liquid water

content that can be made. The visible image of 0930 Z in figure 5.1 shows a bright patch of cloud south of Porto Santo which matches the broad area of increased cloud liquid water in figure 5.12. Likewise, the cloudless area west and southwest of Porto Santo seen in the visible image of 1830 Z (figure 5.3) also compares to the lack of cloud liquid water seen in the image in figure 5.14.

5.2.2 Vertical distribution of liquid water in specified clouds

The liquid water content is computed directly from an average liquid water path and the radar derived cloud thickness using

$$w = \frac{W}{\Delta z} \quad (5.1)$$

Average liquid water path is computed from the 25 data points surrounding the site (see figure 3.6).

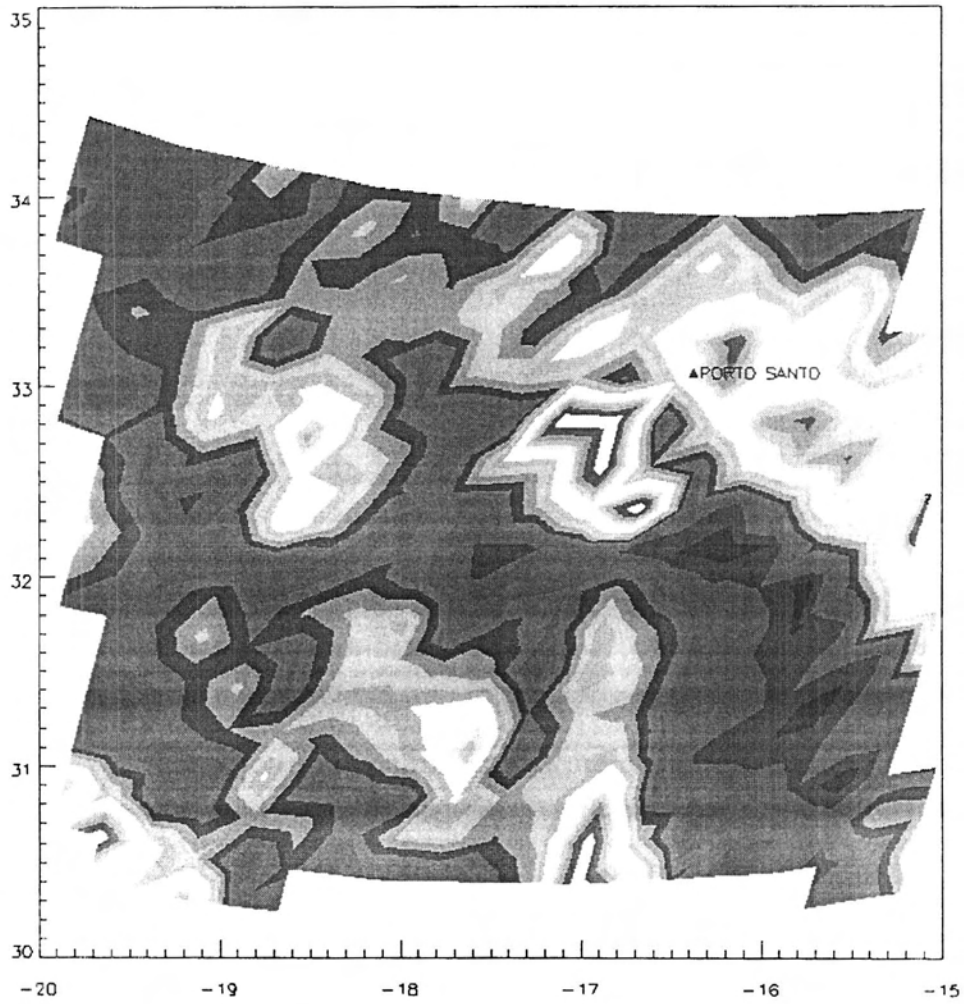
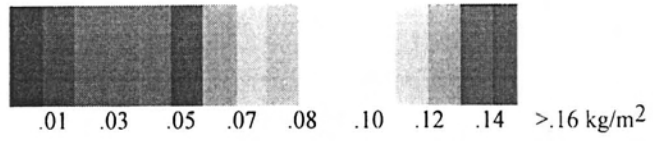


Figure 5.11: LWP retrieved using SSM/I microwave brightness temperatures at 0700Z.

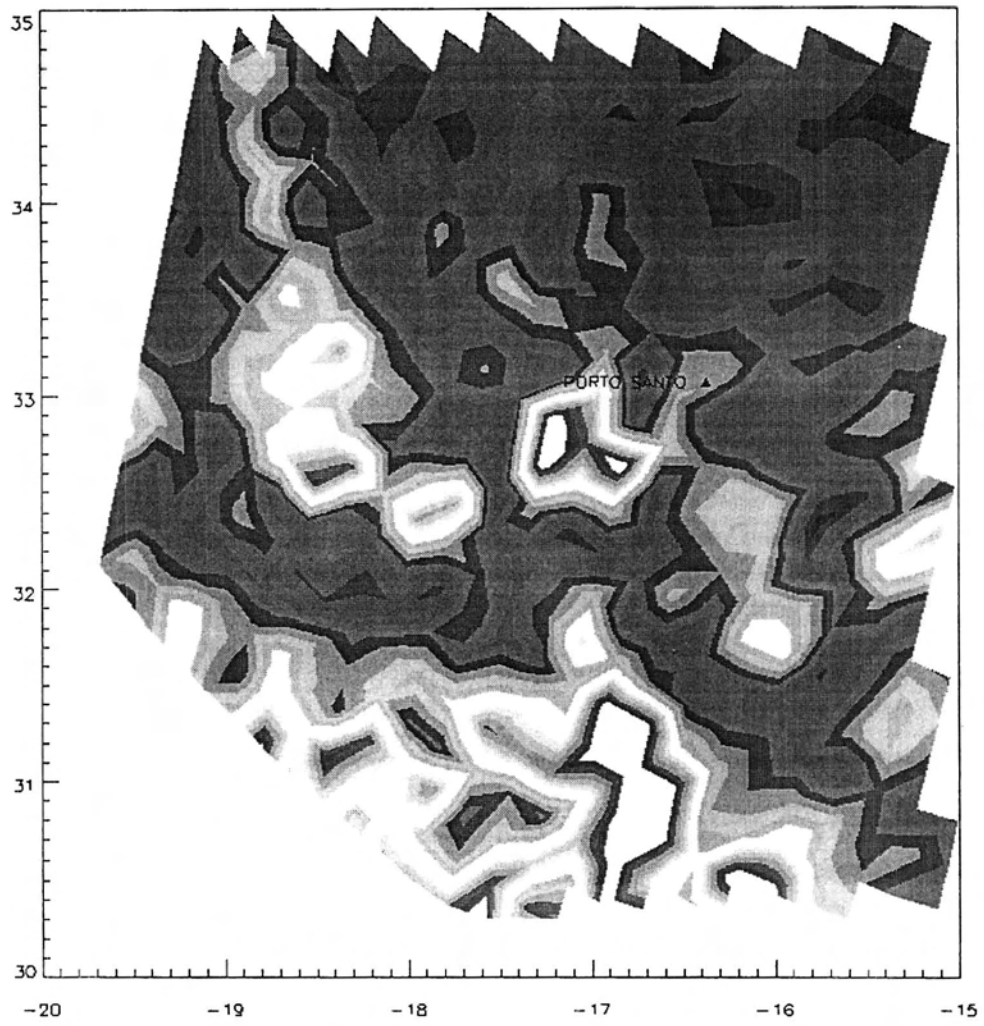
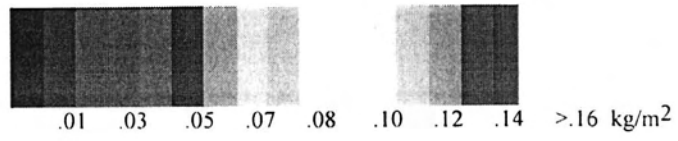


Figure 5.12: LWP retrieved using SSM/I microwave brightness temperatures at 1000Z.

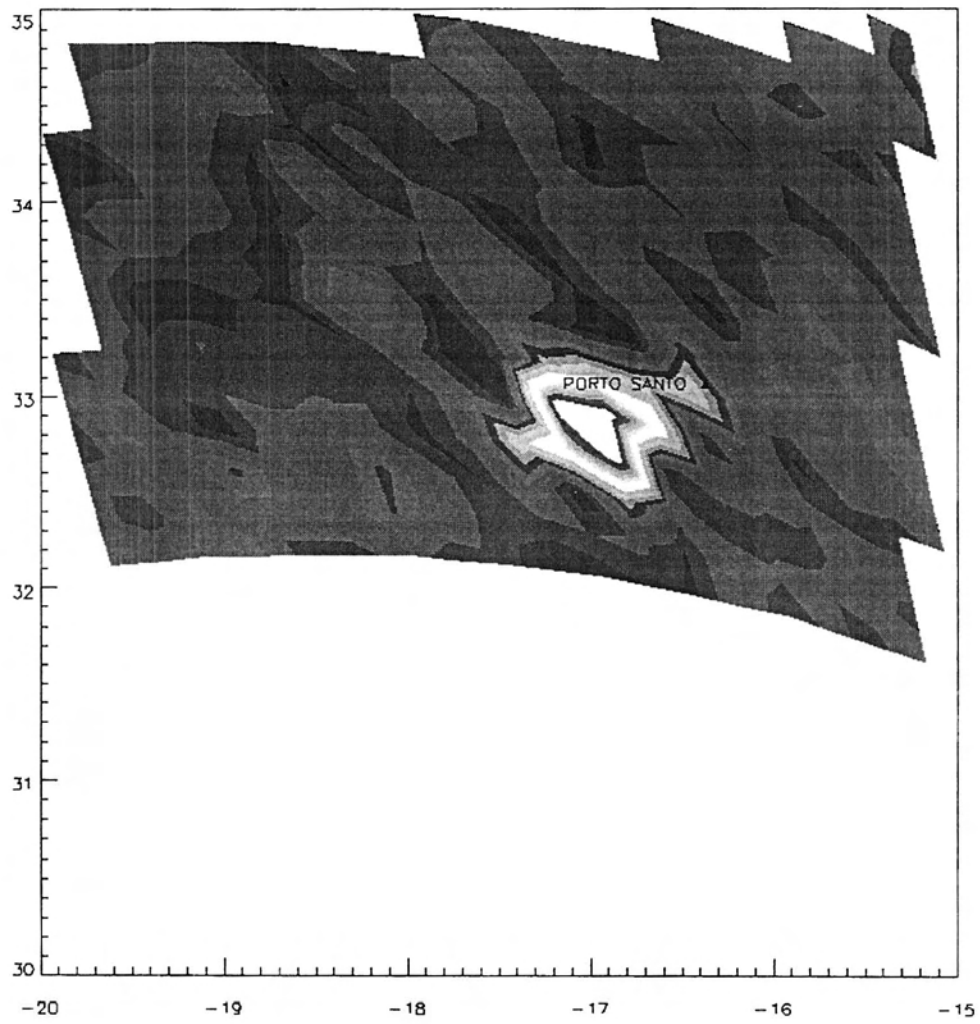
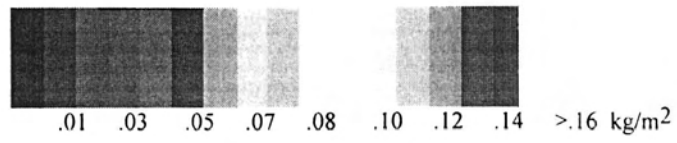


Figure 5.13: LWP retrieved using SSM/I microwave brightness temperatures at 1800Z.

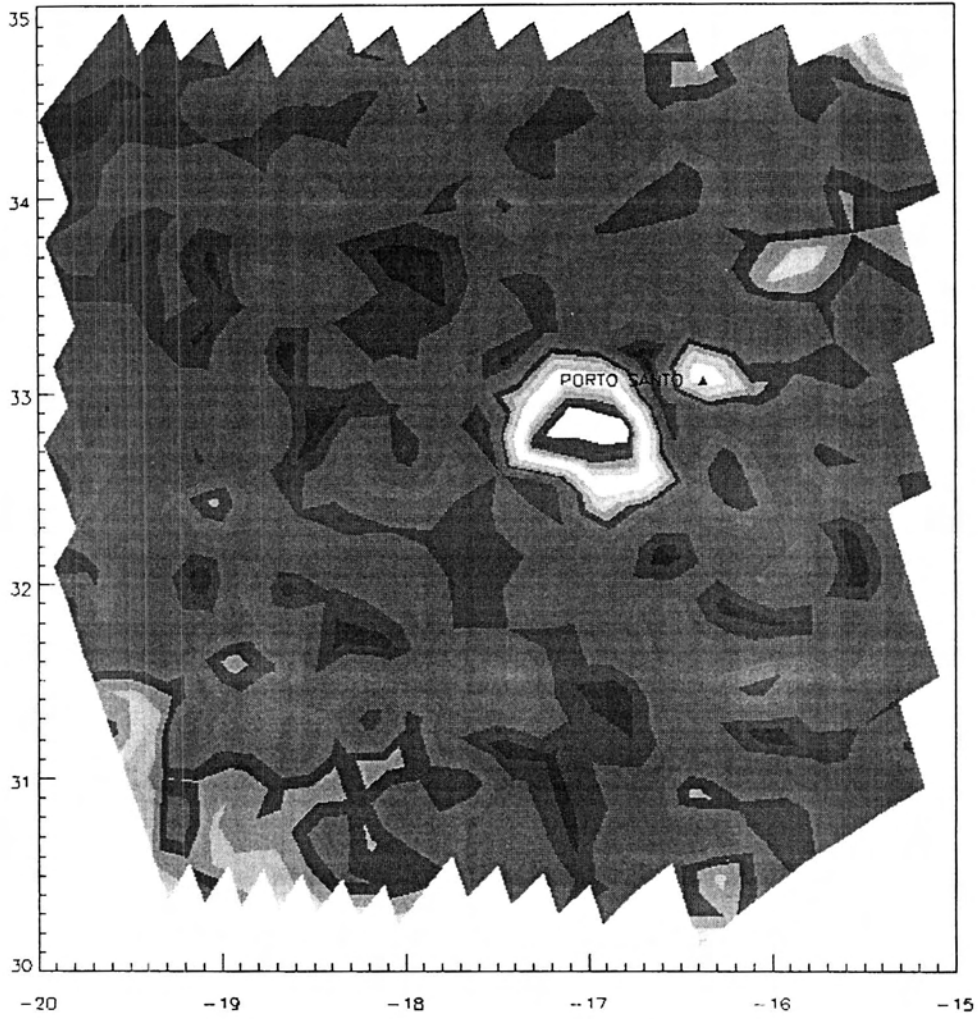
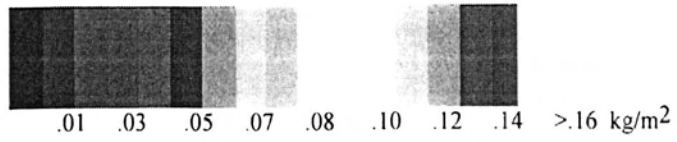


Figure 5.14: LWP retrieved using SSM/I microwave brightness temperatures at 2100Z.

Figure 5.15 shows the location of the cloud layer throughout the day along with the measured liquid water path and computed liquid water content of the cloud at each analysis time. The liquid water path decreases and then increases slightly with time, but with time the cloud gets thinner and this results in higher liquid water contents.

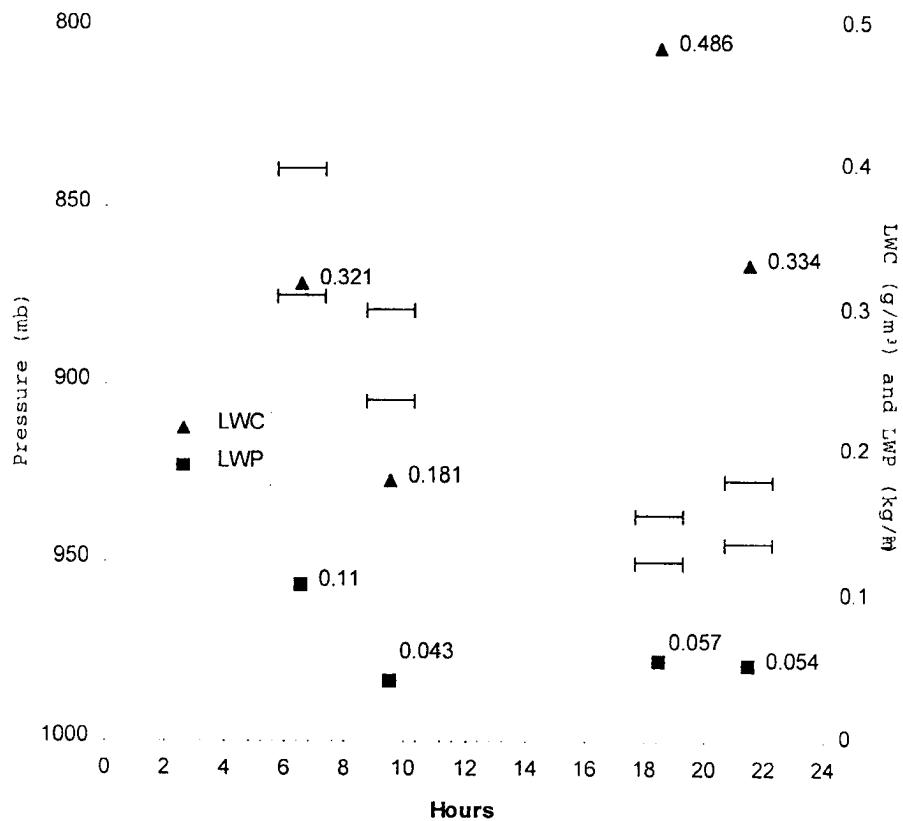


Figure 5.15: Location of cloud base and top with LWC and LWP.

5.3 Heating and cooling rates

Figure 5.16 shows the vertical profile of heating and cooling computed for the cloud layers at 0630Z, 0930Z, 1830Z and 2130Z. The location of the cloud is evident by the strong cooling at the top of the cloud and the warming at the base. There is little

warming due the low height of the cloud and the small difference between the temperature of the cloud base and surface. The maximum cooling occurs at 0630Z and is the result of the higher height of the cloud top coupled with high water content compared to the cloud at later times. As was seen in section 2.2.5, as a cloud layer descends, cooling at the top decreases with the increase in cloud top temperature. However, the decrease in cooling is countered by the increase in liquid water content which increases cooling, so the peak cooling rate actually increases from 0930Z to 2130Z.

Figure 5.17 shows the heating and cooling rates integrated over 200 mb layers for the four time periods. Net cooling exists throughout the atmosphere with significant cooling in the lowest layer where the cloud exists. The cooling rates range from -0.52 K/day to -0.45 K/day; therefore infrared cooling from this cloud layer is considered to influence the diabatic heating and cooling of the atmosphere.

The heating profiles were computed again without the cloud to measure the impact of the cloud on the vertical heating and cooling profile. The average heating rates are shown in figure 5.18. There is still cooling throughout with a larger amount of cooling in the lowest 200 mb. Figure 5.19 shows the difference between the cooling with and without clouds for each level (cloudy-clear). As expected, the cloud has the largest impact in the lowest level and a negligible cooling to the atmosphere above the cloud up to 200 mb. As the observed cloud decreases in altitude, the effect on the atmospheric cooling increases in the lowest layer and decreases in the atmosphere above the cloud. Figure 5.20 shows the percent change from cloudy to clear conditions. The change is all confined to the lowest layer.

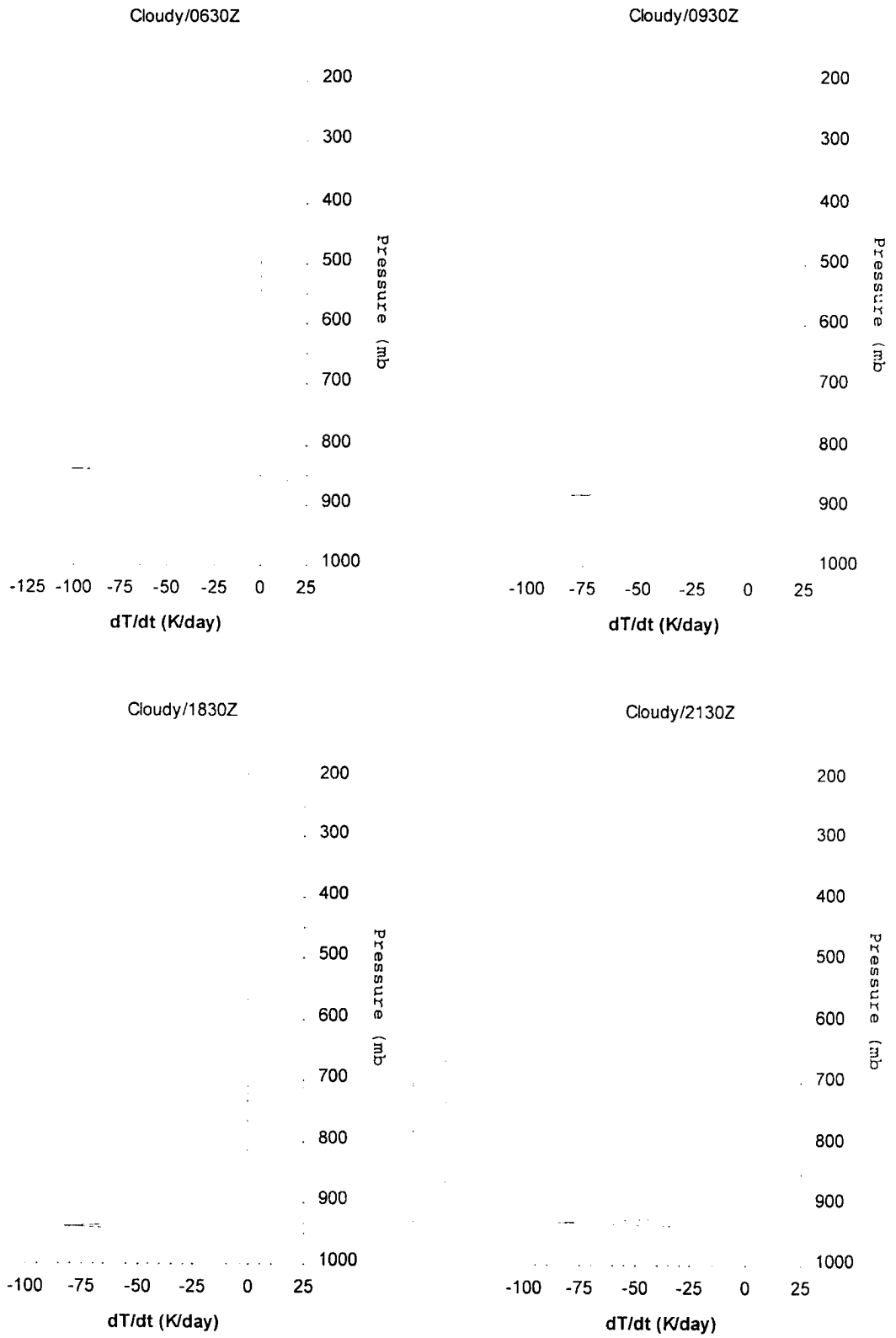


Figure 5.16: Heating and cooling profiles for each analysis time.

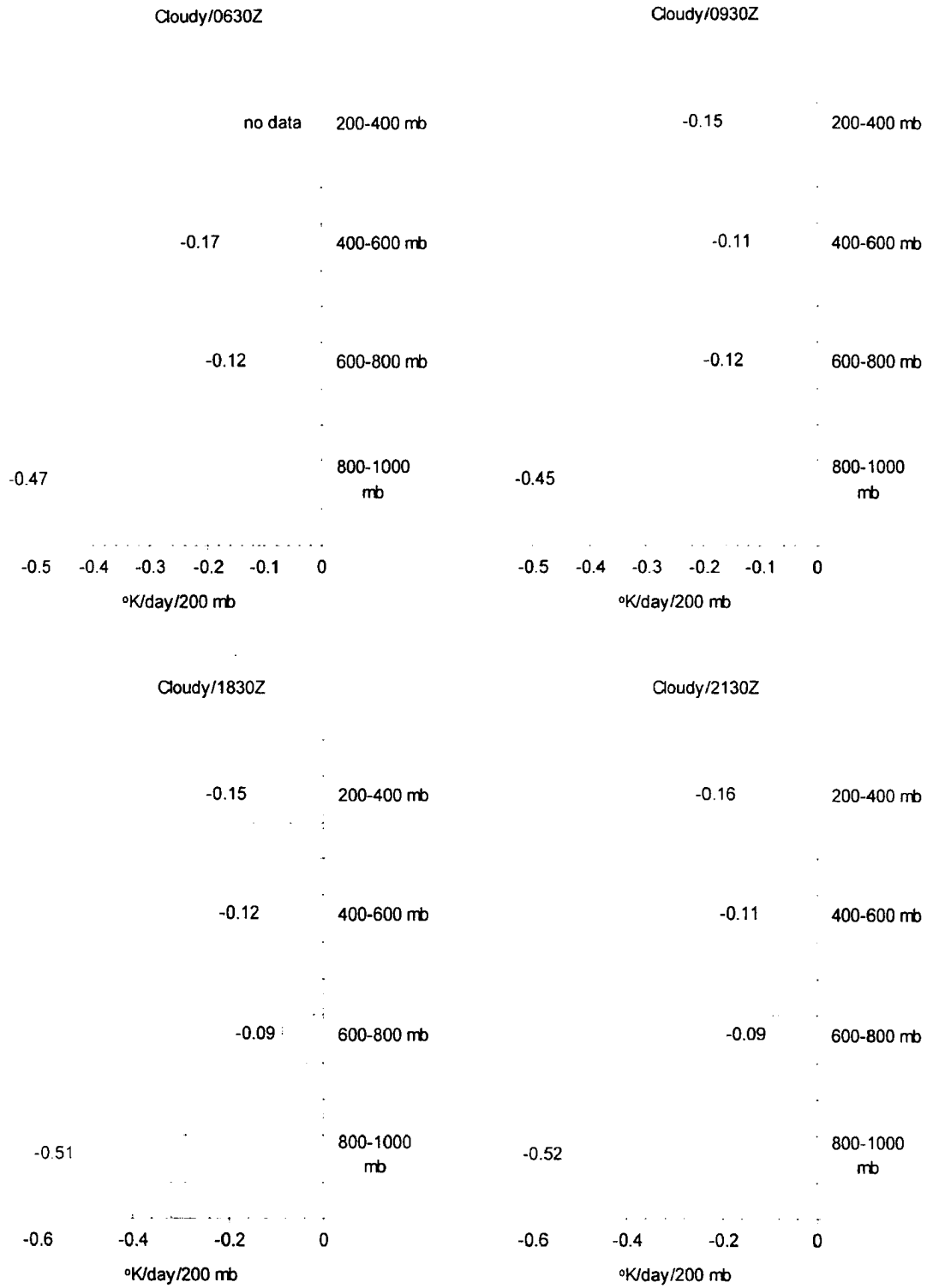


Figure 5.17: Average heating rate per 200 mb layers.

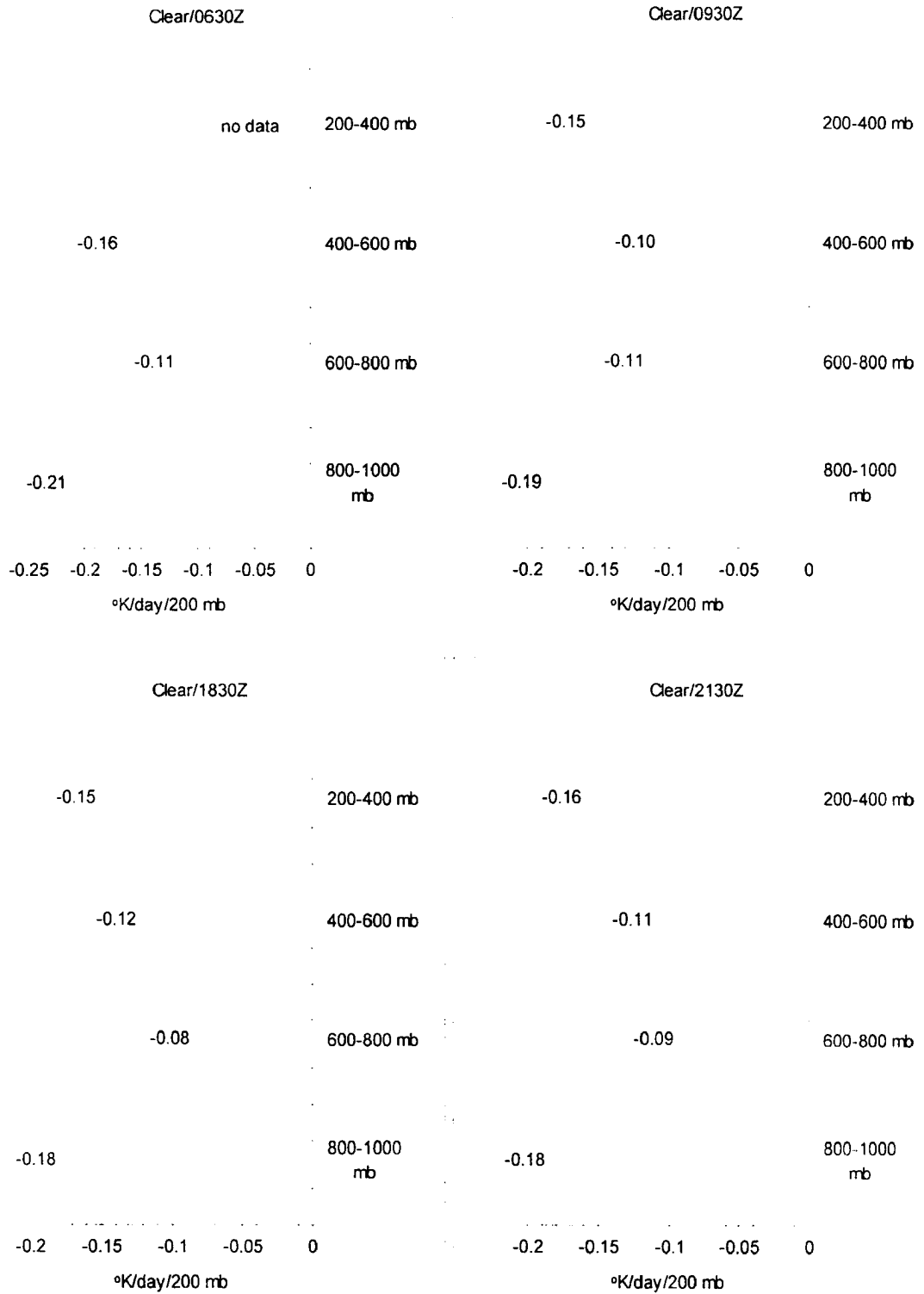


Figure 5.18: Average heating rate per 200 mb layers without the cloud layer.

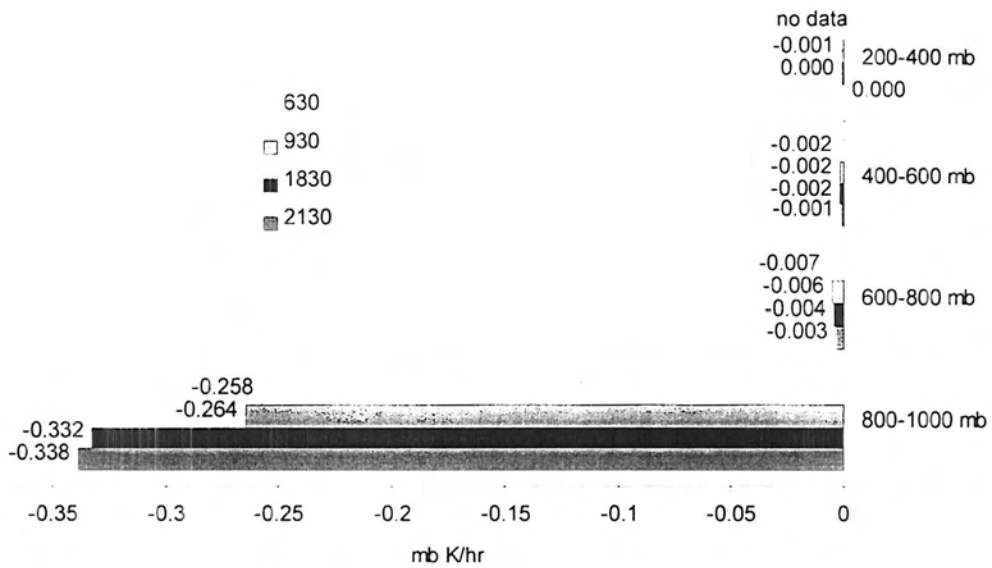


Figure 5.19: Change in heating rate due to cloud (cloud-clear).

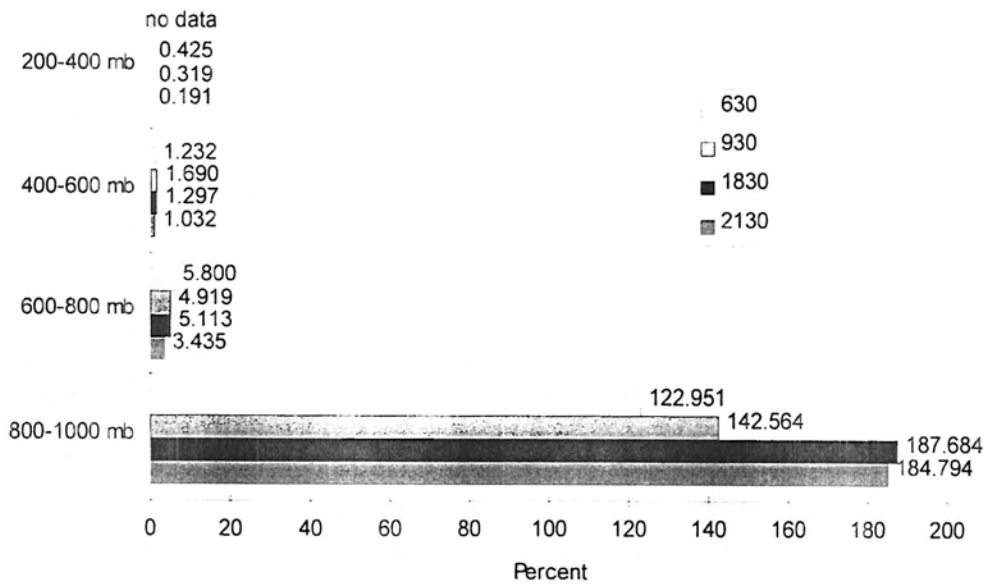


Figure 5.20: Percent change in heating rate due to cloud.

CHAPTER 6

CONCLUSION

This research investigates the modeled infrared radiative heating and cooling profile in the atmosphere that occurs with an observed cloud scenario. Two different situations were considered; one tropical with multiple layers of cumuliform clouds and one sub-tropical with a single layer of stratiform cloud. Observations of the clouds were made primarily using satellite imagery from the geostationary satellites, GMS and Meteosat-3, and the polar orbiting DMSP satellite. Case selection was limited to times when coincident data were available and also to a marine environment for the measurements of integrated cloud liquid water. In Case I, the cloud scene directly over three rawinsonde launch sites was quantified and used in the radiative computations. There was only one site available for Case II, but more DMSP data was available, so the change in the cloud was tracked over one day.

Satellite infrared and visible imagery was used to determine cloud type and the height of the highest cloud tops. The vertical cloud profile was determined from either rawinsonde humidity profiles or millimeter wavelength Doppler cloud radar. The rawinsonde method identified cloud bases/tops based on an increase/decrease in relative humidity. Since there was no way of knowing if the sounding balloon ascended directly over the site and there was no knowledge of the homogeneity of the cloud, the location of the cloud layers in Case I is prone to error. Availability of radar and ceilometer data in addition to rawinsonde humidity profiles allowed for a more accurate measurement of the cloud base and top in Case II.

An integrated cloud liquid water path over the sites in each case was obtained from microwave brightness temperatures measured by the SSM/I instrument. The liquid water path was then divided among the cloud layers based on the temperature and a presumed cloud type. This vertical profile of liquid water content and cloud were then input into a simple radiative transfer program to compute the downwelling and upwelling fluxes and the radiative heating and cooling rates.

As expected, the clouds impact the atmosphere through cooling at the top of the cloud and warming at the base. In the multiple layer clouds, cooling at the tops of lower level clouds was reduced because of the increase downwelling flux from clouds above.

The heating and cooling rates were integrated over 200 mb layers to look at the overall impact on the atmosphere. A heating/cooling rate of ± 2 °K/day/200 mb indicates a threshold at which radiative effects are realized when compared to the magnitudes of latent and sensible heating. In Case I, the cooling in the 200-400 mb layer was less than -0.258 °K/day/200 mb at all sites and in some cases was less than -0.5 °K/day/200 mb. Most of the cloud layers had tops in the 200-400 mb layer and bases in the 400-600 mb and 600-800 mb layers. The warming from the bases of these clouds resulted in a slight warming in these layers, but not on a large enough scale to have an effect. The more noticeable effect is the vertical heating rate gradient between the layers which leads impacts atmospheric dynamics, in particular it causes destabilization of the cloud layers.

In Case II, a single cloud was located entirely in the lowest layer (800 mb -1000 mb). The cloud affected this layer through radiative cooling of between -0.47 °K/day/200

mb and -0.51 °K/day/200 mb. Cooling effects on the atmosphere above the cloud were negligible. Because of the low level of the cloud, warming at the base was small.

Cloud forcing was investigated by looking at the vertical heating profile with and without the observed cloud layers. For the multiple cloud layers, the clouds were found to increase the cooling in the upper levels of the atmosphere and decrease the cooling in the lower levels. The largest change in the vertical heating and cooling rates occurred in the middle layers of the atmosphere. Cloud forcing study of the single cloud in Case II, only revealed as increase in cooling in the lowest layer of the atmosphere because the entire cloud was in the lowest 200 mb region and the warming at the base of the cloud was small.

The largest sources of error and room for improvement are in determination of the vertical cloud profile and the distribution of the liquid water in those layers. Future lidar and millimeter wavelength imagers flown on satellites will aid in determining the location of cloud bases and tops and will increase the regions over which these studies can be performed by eliminating or reducing the reliance on ground instruments. The distribution of liquid water among the layers could be improved by capturing the true microphysics of the cloud to represent the actual vertical distribution and content of water and ice in the cloud layers.

Finally, the measurement of cloud liquid water over the tropics is a source of error due to the rainfall and cirrus associated with tropical convection. Measurement of cloud liquid water over land would allow for investigation of cloud scenes over land where

more corroborative ground data such as radar and sounding data is available to better model the true vertical profile of cloud and cloud water.

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